

Exact QM solution in one-electron atoms.

McMurry Ch 7.1 and 7.2. SP Ch 2.1, Foot Ch 2.

Exercises: 12 - 15. (H4 and H5)

$$\text{SE: } \left(-\frac{\hbar^2}{2\mu} \nabla^2 - \frac{Ze^2}{4\pi\epsilon_0 r} \right) \Psi = E\Psi$$

$$E_n = -R_M \cdot \frac{Z^2}{n^2}$$

$$\Psi_{n,\ell,m}(r, \theta, \varphi) = R_{n,\ell}(r) \cdot Y_{\ell,m}(\theta, \varphi)$$

$$\int_0^\infty \int_0^\pi \int_0^{2\pi} |R_{n,\ell}(r) \cdot Y_{\ell,m}(\theta, \varphi)|^2 r^2 \sin\theta dr d\theta d\varphi = 1$$

$$R(\rho) = \rho^\ell \cdot e^{-\rho} \cdot c_0 \cdot \left(1 + \sum_{k=0}^{n-\ell-1} \frac{c_k}{c_0} \cdot \rho^k \right), \quad \rho = \frac{Z}{na_0} \cdot r$$

$c_0 \neq 0$ determined from the normalization condition.

$$c_{k+1} = -2c_k \frac{n - (\ell + k + 1)}{(\ell + k + 2)(\ell + k + 1) - \ell(\ell + 1)}, \quad k = 0, 1, \dots, n - \ell - 1$$

$$\langle r_{n\ell} \rangle = \int R^*(r) \cdot r \cdot R(r) \cdot r^2 dr = \frac{1}{2Z} [3n^2 - \ell(\ell + 1)]$$

One-electron atoms in spherical coordinates

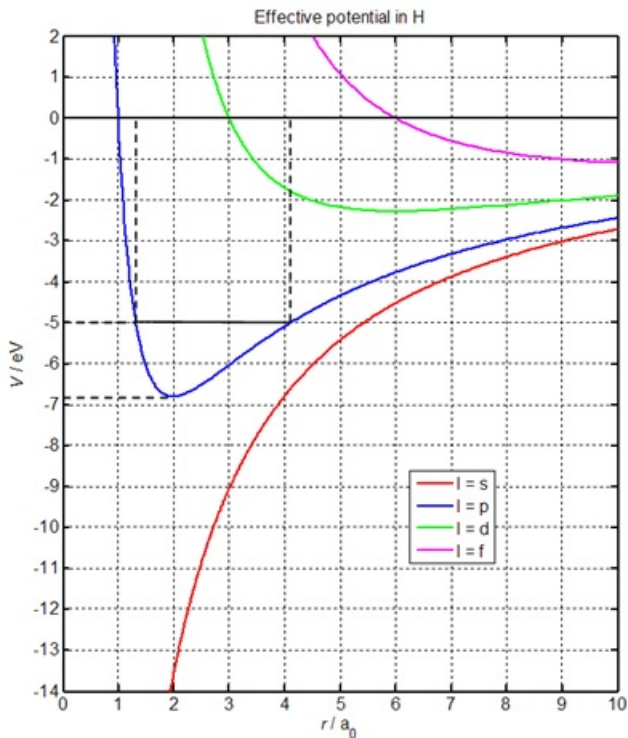
$$L_x = i\hbar\left(\sin\varphi\frac{\partial}{\partial\theta} + \frac{\cos\varphi}{\tan\theta}\frac{\partial}{\partial\varphi}\right),$$

$$L_y = i\hbar\left(-\cos\varphi\frac{\partial}{\partial\theta} + \frac{\sin\varphi}{\tan\theta}\frac{\partial}{\partial\varphi}\right),$$

$$L_z = -i\hbar\frac{\partial}{\partial\varphi}$$

$$L^2 = -\hbar^2\left(\frac{\partial^2}{\partial\theta^2} + \frac{1}{\tan\theta}\frac{\partial}{\partial\theta} + \frac{1}{\sin^2\theta}\frac{\partial^2}{\partial\varphi^2}\right)$$

$$\begin{aligned}H &= -\frac{\hbar^2}{2\mu}\left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}\right) - \frac{Ze^2}{4\pi\epsilon_0\cdot r} \\&= -\frac{\hbar^2}{2\mu}\left[\frac{1}{r}\frac{\partial^2}{\partial r^2}r + \frac{1}{r^2}\left(\frac{\partial^2}{\partial\theta^2} + \frac{1}{\tan\theta}\frac{\partial}{\partial\theta} + \frac{1}{\sin^2\theta}\frac{\partial^2}{\partial\varphi^2}\right)\right] - \frac{Ze^2}{4\pi\epsilon_0\cdot r} \\&= -\frac{\hbar^2}{2\mu}\cdot\frac{1}{r}\frac{\partial^2}{\partial r^2}r + \frac{L^2}{2\mu\cdot r^2} - \frac{Ze^2}{4\pi\epsilon_0\cdot r}\end{aligned}$$



$$V_{eff} = \frac{\hbar^2 \ell \cdot (\ell + 1)}{2\mu \cdot r^2} - \frac{Ze^2}{4\pi\epsilon_0 \cdot r}$$

In atomic physics the value of ℓ is always given as a letter:
 $\ell = 0 \rightarrow s$, $1 \rightarrow p$, $2 \rightarrow d$, $3 \rightarrow f$, $4 \rightarrow g$, h , i , (j) , k , l , ...

A p electron with an energy of -5 eV is classically constrained to oscillate between $1.4 < r < 4 a_0$ whereas with an energy of about -6.8 eV it can only move in a circular orbit with a radius of $2a_0$.

**Note, only s-electrons can move close to the nucleus.
 Very important later on!!**

Series solution of the radial differential equation.

$$f'' - 2f' - \left(\frac{\ell(\ell+1)}{\rho^2} - \frac{2Z}{a_0\kappa\rho} \right) f = 0 \quad (8)$$

Ansatz:

$$f(\rho) = \rho^s \cdot \sum_{k=0}^{\infty} c_k \rho^k \quad c_0 \neq 0.$$

Calculate the derivatives and collect terms with similar powers in (8)

$$\sum_{k=0}^{\infty} c_k [(s+k)(s+k-1) - \ell(\ell+1)] \rho^{s+k-2} + \sum_{k=0}^{\infty} c_k \left[\frac{2Z}{a_0\kappa} - 2(s+k) \right] \rho^{s+k-1} = 0$$

The lowest order term arise from the first sum with $k = 0$. If we write that term separately the first sum will start with $k = 1$ and both sums will be in powers of $s + k - 1$. If we change summation index in the first sum from k to $k + 1$ we can write everything as a single summation:

$$c_0 \cdot [s(s-1) - \ell(\ell+1)] \rho^{s-2} + \sum_{k=0}^{\infty} \{ c_{k+1} \cdot [(s+k+1)(s+k) - \ell(\ell+1)] + 2c_k \cdot \left[\frac{Z}{a_0\kappa} - (s+k) \right] \} \rho^{s+k-1} = 0$$

If the polynomial is zero for all values of ρ then each coefficient must be zero individually.

Since we assumed $c_0 \neq 0$ the first term gives $s = \ell + 1$ or $s = -\ell$. The last possibility must be ignored since it leads to a non-normalizable function $P(\rho)$ when $\rho \rightarrow 0$.

To make everything within the curly brackets equal to zero in the remaining summation the c -coefficients must satisfy the recursion relation (using $s = \ell + 1$)

$$c_{k+1} = -2c_k \frac{(Z/a_0\kappa) - (\ell + k + 1)}{(\ell + k + 2)(\ell + k + 1) - \ell(\ell + 1)}$$

For large k the ratio of the coefficients will be $c_{k+1}/c_k = 2k/k^2 = 2/k$. This ratio is the same as that between successive terms in the expansion of

$$e^{2\rho} = 1 + \frac{2\rho}{1!} + \frac{2^2}{2!} \rho^2 + \dots + \frac{2^{k-1}}{(k-1)!} \rho^{k-1} + \frac{2^k}{k!} \rho^k + \dots$$

Hence, for large ρ the two functions behave similar and $P(\rho) \approx e^{-\rho} \cdot e^{2\rho} \rightarrow \infty$ when $\rho \rightarrow \infty$.

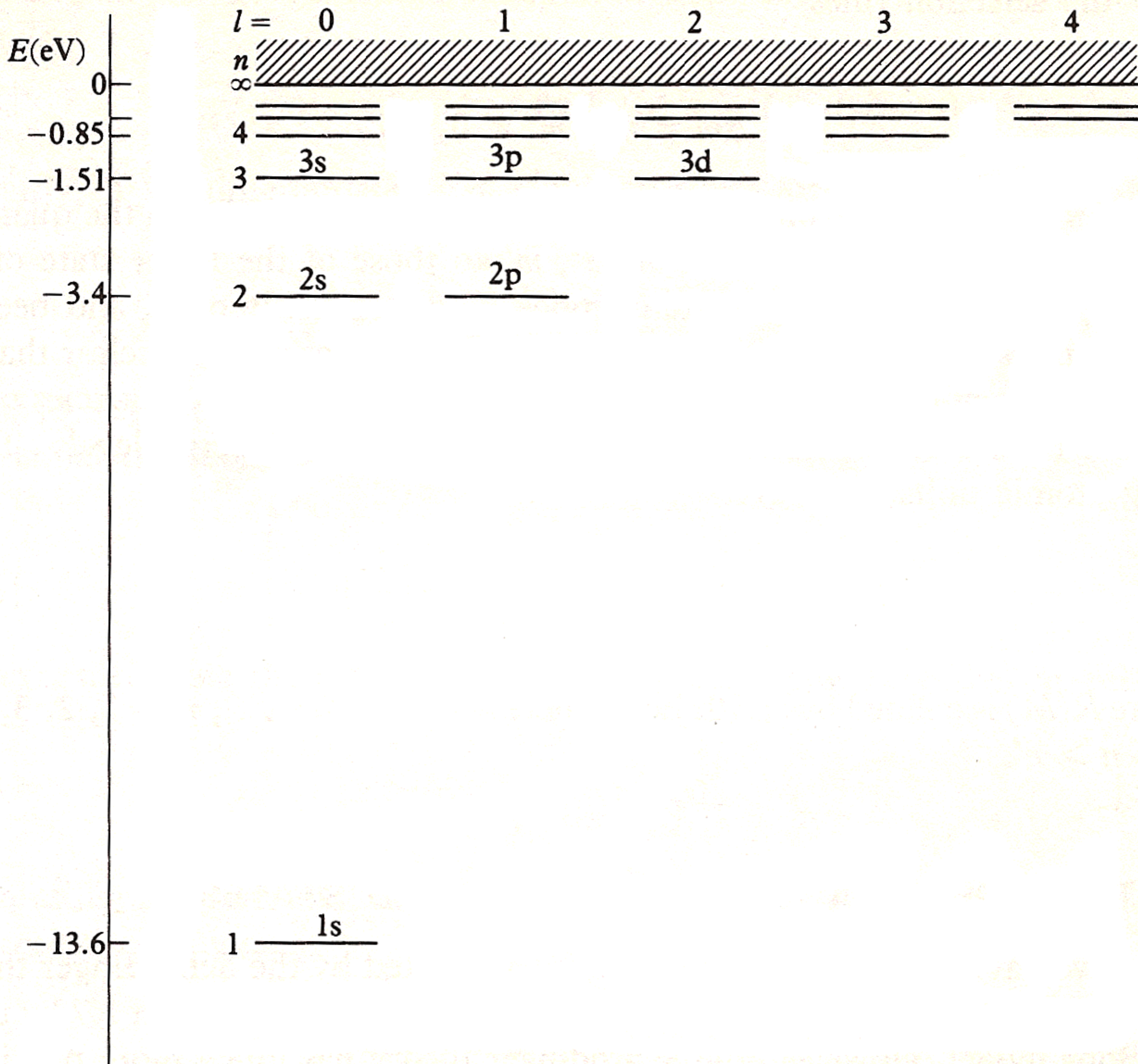
Thus it is not normalizable and the series must be truncated, i.e. for some

$$k = N = 0, 1, 2, \dots \quad (Z/a_0\kappa) - (\ell + N + 1) = 0.$$

Introducing the main quantum number $n = \ell + N + 1 = 1, 2, 3, \dots$ we obtain finally $\kappa = Z/a_0 n$ and hence a quantized energy:

$$E_n = -\frac{\hbar^2}{2\mu a_0^2} \cdot \frac{Z^2}{n^2}.$$

Thus the reasonable demand that the wavefunction is normalizable leads to a quantized energy!!



3.2 Energy-level diagram for atomic hydrogen.

Table 2.2 Radial hydrogenic wavefunctions $R_{n,l}$ in terms of the variable $\rho = Zr/(na_0)$, which gives a scaling that varies with n . The Bohr radius a_0 is defined in eqn 1.40.

$$R_{1,0} = \left(\frac{Z}{a_0}\right)^{3/2} 2 e^{-\rho}$$

$$R_{2,0} = \left(\frac{Z}{2a_0}\right)^{3/2} 2(1-\rho) e^{-\rho}$$

$$R_{2,1} = \left(\frac{Z}{2a_0}\right)^{3/2} \frac{2}{\sqrt{3}} \rho e^{-\rho}$$

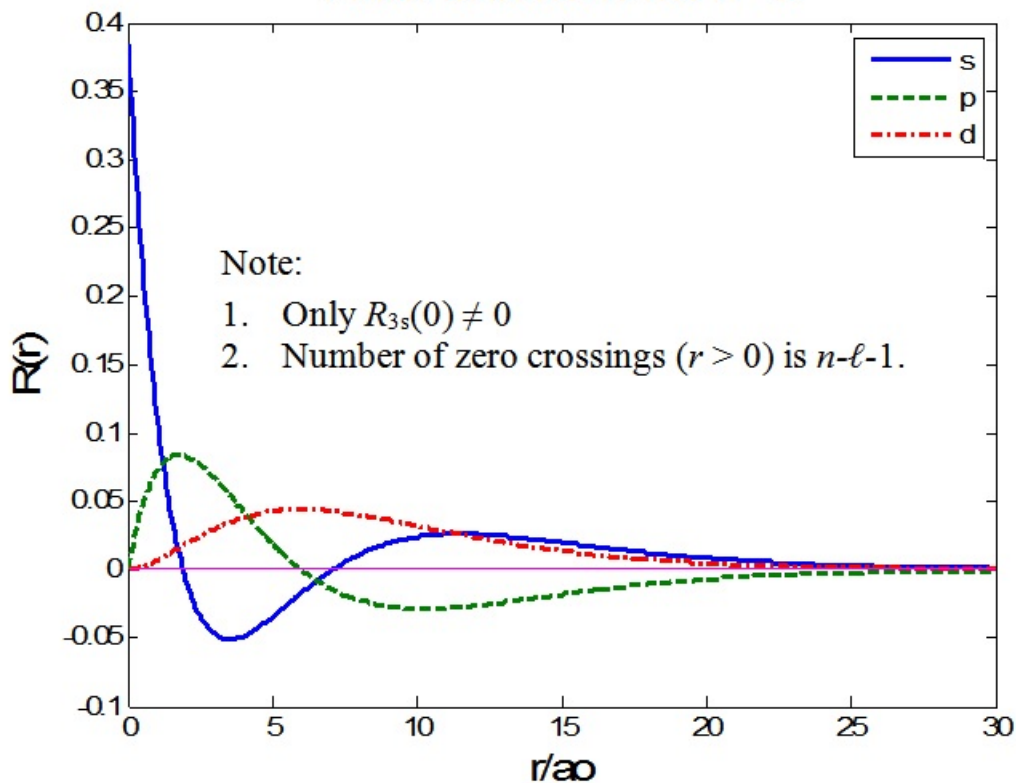
$$R_{3,0} = \left(\frac{Z}{3a_0}\right)^{3/2} 2 \left(1 - 2\rho + \frac{2}{3}\rho^2\right) e^{-\rho}$$

$$R_{3,1} = \left(\frac{Z}{3a_0}\right)^{3/2} \frac{4\sqrt{2}}{3} \rho \left(1 - \frac{1}{2}\rho\right) e^{-\rho}$$

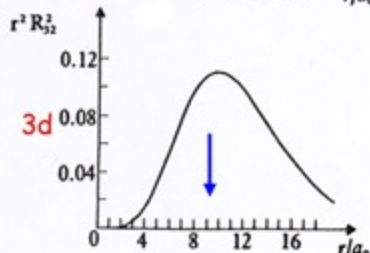
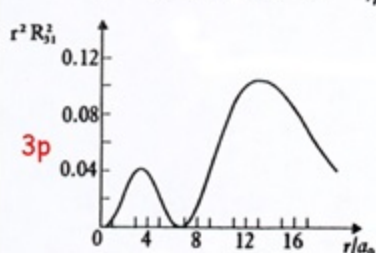
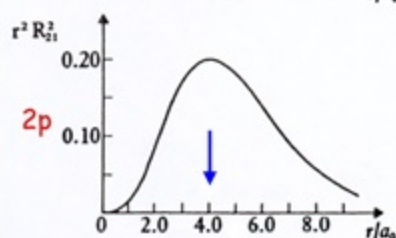
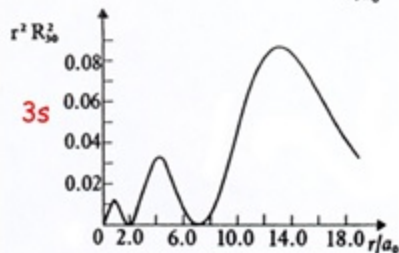
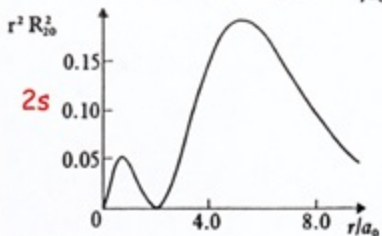
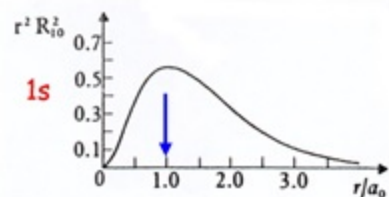
$$R_{3,2} = \left(\frac{Z}{3a_0}\right)^{3/2} \frac{2\sqrt{2}}{3\sqrt{5}} \rho^2 e^{-\rho}$$

Normalisation: $\int_0^\infty R_{n,l}^2 r^2 dr = 1$

Radial functions in H for $n=3$



Radial probability distribution in H.



For 1s, 2p, 3d,... i.e.

$$n, \ell = n - 1$$

the most probable distance r_n is given by:

$$r_n = a_0 \frac{n^2}{Z}$$

i.e. the same result as in the Bohr model!

Number of minima:

$$N = n - \ell - 1$$

Summary one-electron atoms.

Bohr:

$$r_n = a_0 \cdot \frac{n^2}{Z}, \quad a_0 = 0.529 \text{ \AA}.$$

$$v_n = \alpha \cdot c \cdot \frac{Z}{n}, \quad \alpha = \frac{1}{137}$$

$$E_n = -R_M \cdot \frac{Z^2}{n^2}, \quad R_M = \frac{m}{M+m} \cdot 109737 \text{ cm}^{-1}$$

Quantum mechanics

$$\left(-\frac{\hbar^2}{2\mu} \nabla^2 - \frac{Ze^2}{4\pi\epsilon_0 r}\right)\Psi = E\Psi$$

$$E_n = -R_M \cdot \frac{Z^2}{n^2} \quad (E \text{ in cm}^{-1})$$

Only for the Coulomb potential is E independent of ℓ .

$$n = 1, 2, 3, \dots \quad \ell = 0, 1, 2, \dots, (n-1) \quad m = -\ell, -\ell+1, \dots, \ell$$

$$\Psi_{n,\ell,m}(r, \theta, \varphi) = R_{n,\ell}(r)Y_{\ell,m}(\theta, \varphi)$$

$$R(\rho) = \rho^\ell \cdot e^{-\rho} \cdot c_0 \cdot \left(1 + \sum_{k=1}^{n-\ell-1} \frac{c_k}{c_0} \cdot \rho^k\right), \quad \rho = \frac{Z}{a_0 n} \cdot r$$

$c_0 \neq 0$ determined from the normalization condition.

$R(\rho)$ has $n - \ell - 1$ zeros for $r > 0$

$$\int_0^\infty \int_0^\pi \int_0^{2\pi} |R_{n,\ell}(r) \cdot Y_{\ell,m}(\theta, \varphi)|^2 r^2 \sin \theta dr d\theta d\varphi = 1$$

$$\langle r_{n\ell} \rangle = \int R^*(r) \cdot r \cdot R(r) \cdot r^2 dr = \frac{1}{2Z} [3n^2 - \ell(\ell+1)]$$