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Synchrotron radiation and applications of synchrotron radiation

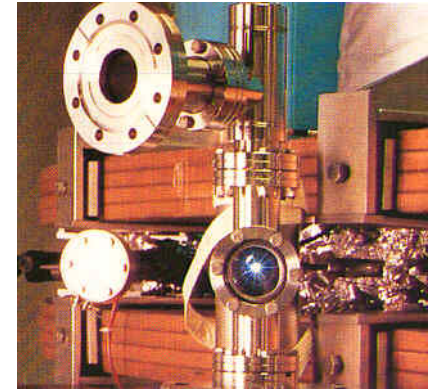
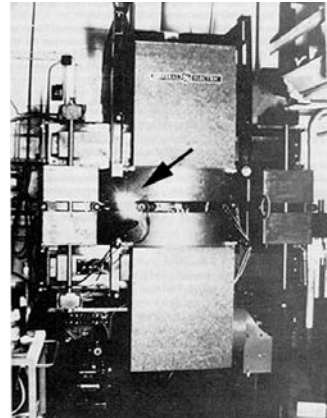
**Joachim Schnadt
Division of Synchrotron Radiation Research
Department of Physics
Lund University**

Synchrotron radiation and synchrotron radiation facilities



What is synchrotron radiation?

Synchrotron radiation was first observed at a synchrotron (at the General Electric Synchrotron Accelerator in 1946). Today synchrotron radiation for use in materials experiments is normally produced in *electron storage rings*.



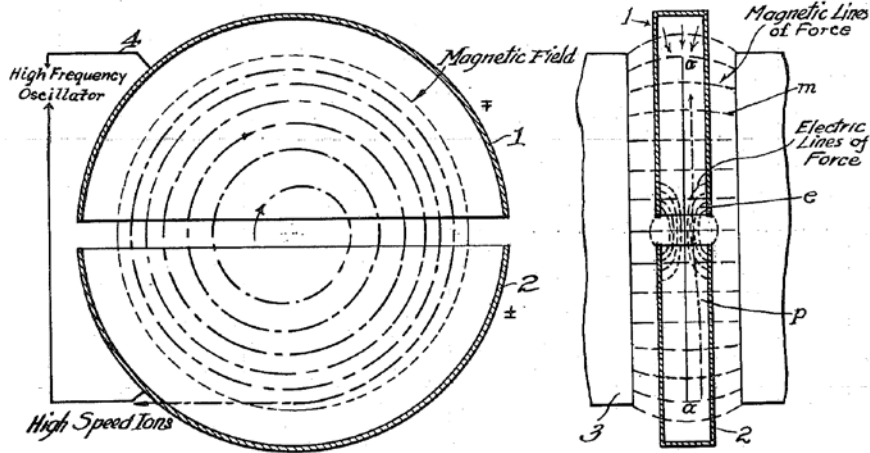
Cyclotron: accelerator with constant magnetic field in electric field-free space and alternating electric field for acceleration, spiral electron trajectory, increasing electron energy

Synchrotron: accelerator with varying magnetic and electric fields, fixed electron trajectory, increasing electron energy

Storage ring: accelerator with fixed electron trajectory and fixed electron energy, electron speed close to speed of light, fields of bending magnets fixed, alternating electric field for compensating energy loss

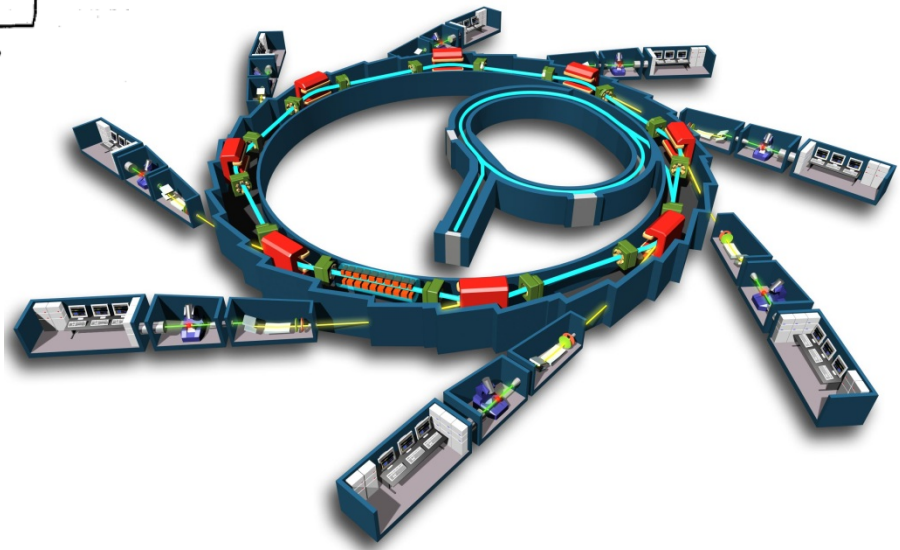


Cyclotron, synchrotron, storage ring



Cyclotron

Synchrotron / storage ring



Synchrotron radiation facilities around the world



www.diamond.ac.uk



Action of a synchrotron radiation source



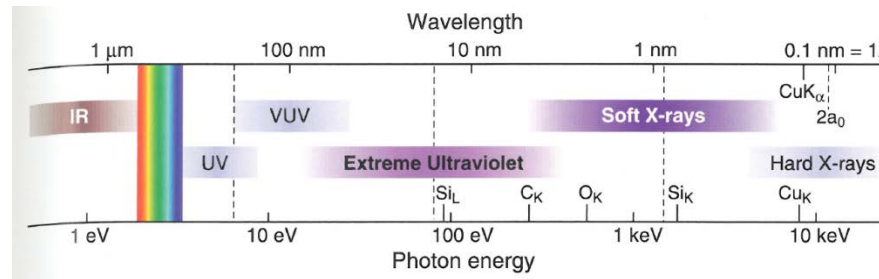
www.isa.au.dk – ASTRID, Aarhus University, Denmark



Properties of synchrotron radiation

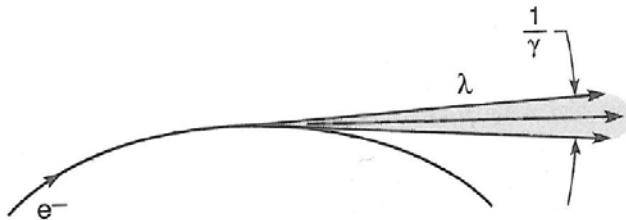
Adapted from Terasawa and Kihara in: H. Saisho and Y. Gohshi (Eds.), *Applications of Synchrotron Radiation to Materials Analysis*, Elsevier, Amsterdam, 1996

(1) A continuous spectrum from the infrared to the X-ray region.



(2) High intensity, owing to the high current electrons accumulated in the storage ring.

(3) Collimation of the emitted radiation in the instantaneous direction of flight of the emitting particles (the angular spread is of the order of 1 mrad).

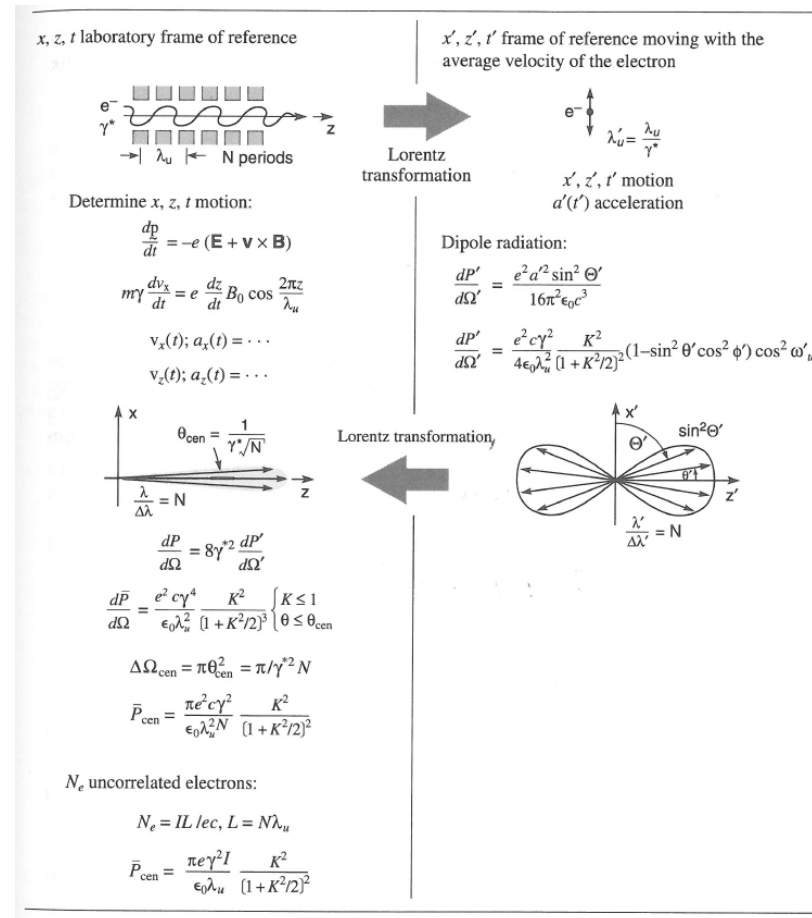


A relativistic speeds (Lorentz contraction factor $\gamma = 1/(1-v^2/c^2)^{1/2}$) electrons emit into the forward direction.



Images: D. Attwood, *Soft X-rays and Extreme Ultraviolet Radiation*, Cambridge University Press, Cambridge, 1999.

Light emission of relativistic electrons



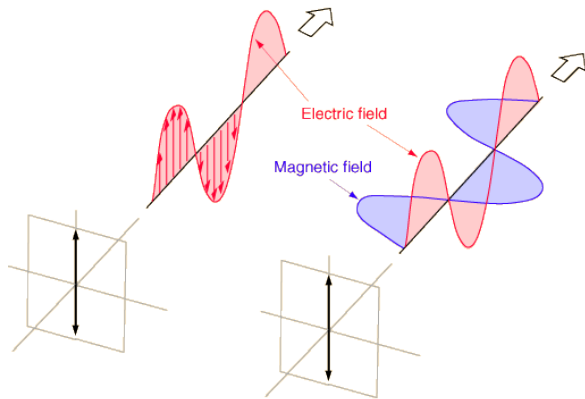
D. Attwood, Soft X-rays and Extreme Ultraviolet Radiation, Cambridge University Press, Cambridge, 1999.



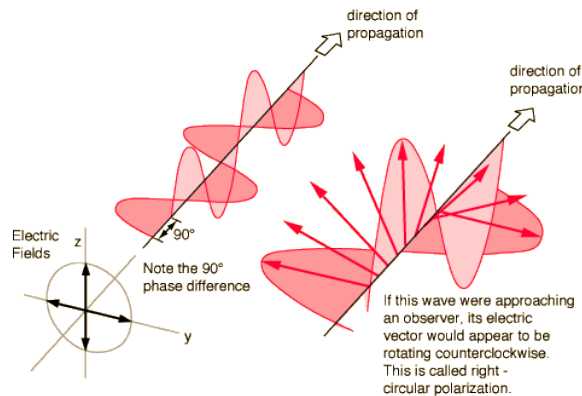
Properties of synchrotron radiation

Adapted from Terasawa and Kihara in: H. Saisho and Y. Gohshi (Eds.), *Applications of Synchrotron Radiation to Materials Analysis*, Elsevier, Amsterdam, 1996

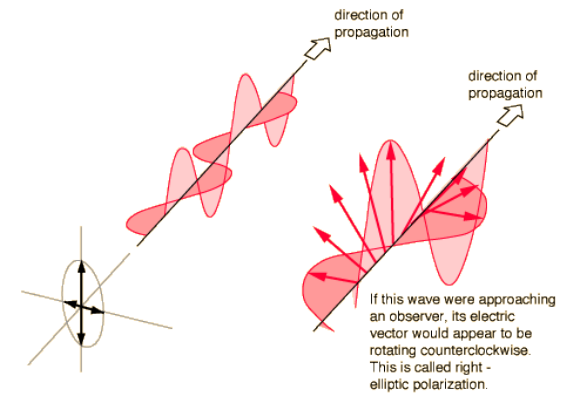
(4) Polarisation control



Linear polarisation



Circular polarisation



Elliptical polarisation

Images: hyperphysics.phy-astr.gsu.edu.



Properties of synchrotron radiation

Adapted from Terasawa and Kihara *in*: H. Saisho and Y. Gohshi (Eds.), *Applications of Synchrotron Radiation to Materials Analysis*, Elsevier, Amsterdam, 1996

- (5) High brilliance of the source, because of the small cross section of the electron beam and the high degree of collimation of the radiation.

Flux = number of photons/(s mm²)

Brilliance = flux/(mrad² 0.1% BW)

0.1% BW denotes a bandwidth of $10^{-3}\nu$ centered around the frequency ν .

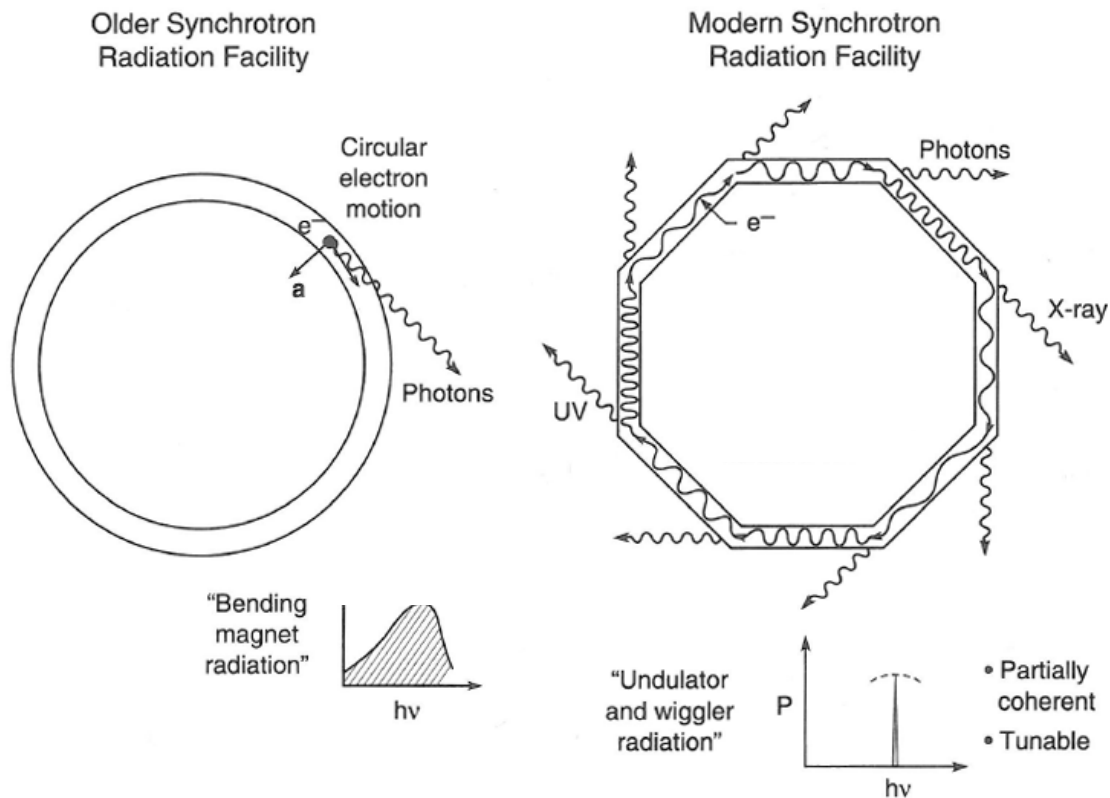
- (6) A time structure with pulse lengths down to 100 ps.

- (7) Absolute calculability of all the properties of the source.

- (8) Cleanliness of the source, since the light emission takes place in an ultra-high vacuum, in contrast to the situation in gas discharge or spark lamps.



Ways of producing synchrotron radiation



Radiation from insertion devices
Radiation from bending magnets (undulators and wigglers) and partly
from bending magnets



Bending magnet radiation

Photon flux per unit angle and 0.1% BW:

$$\left. \frac{d^3 F_B}{d\theta d\psi d\omega/\omega} \right|_{\psi=0} = 1.33 \times 10^{13} E_e^2 (\text{GeV}) I(\text{A}) H_2(E/E_c) \frac{\text{photons/s}}{\text{mrad}^2 \cdot (0.1\% \text{ BW})}$$

H_2 : modified Bessel function (tabulated)

Critical energy: $E_c = \hbar\omega_c = \frac{3e\hbar B\gamma^2}{2m}$

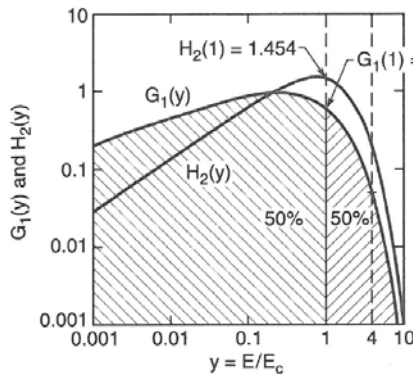
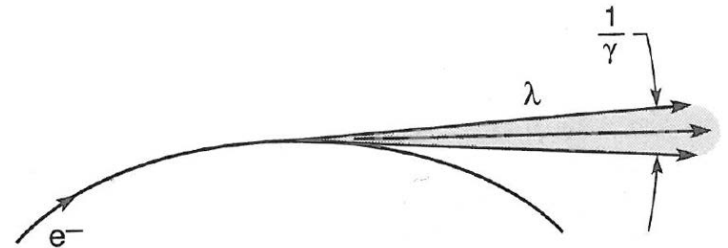
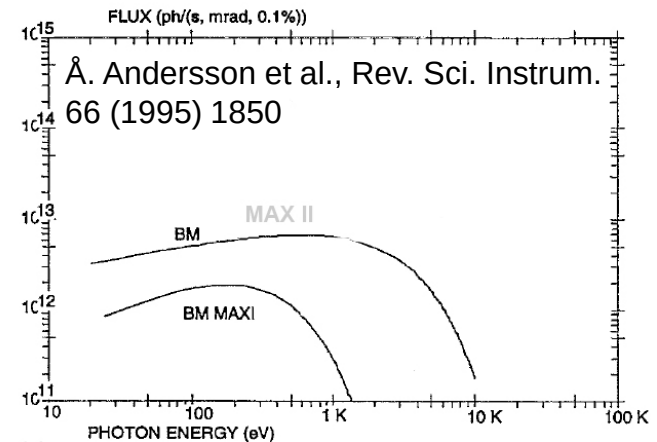


FIGURE 5.7. The functions $H_2(y)$, representing on-axis photon flux from a bending magnet, and $G_1(y)$, representing the vertically integrated photon flux, as functions of photon energy normalized to the critical photon energy. Half the radiated power is in photons of energy greater than E_c , and half in photons of energy less than E_c (following Kim³). Note that for a photon energy of $4E_c$ the photon flux is reduced a factor of about 10 from its value at E_c .



Å. Andersson et al., Rev. Sci. Instrum. 66 (1995) 1850

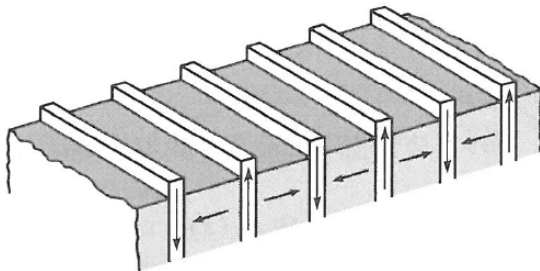
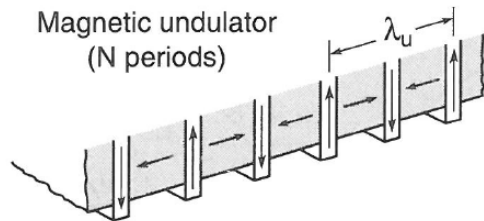


Images: D. Attwood, Soft X-rays and Extreme Ultraviolet Radiation, Cambridge University Press, Cambridge, 1999.

Undulators and wigglers

Undulator equation:

$$\lambda = \frac{\lambda_u}{2\gamma^2} \left(1 + \frac{K^2}{2} + \gamma^2 \theta^2 \right) \quad K \equiv \frac{eB_0\lambda_u}{2\pi mc}$$

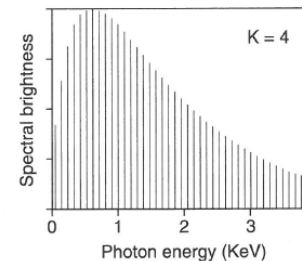
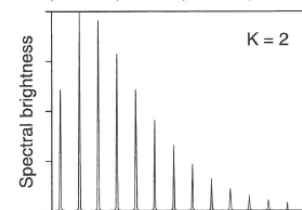
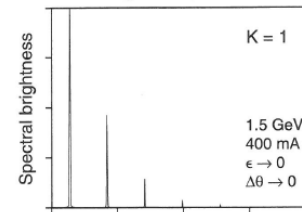


$$\lambda \approx \frac{\lambda_u}{2\gamma^2}$$

$$\theta_{\text{cen}} \approx \frac{1}{\gamma^* \sqrt{N}}$$

$$\left(\frac{\Delta\lambda}{\lambda} \right)_{\text{cen}} = \frac{1}{N}$$

$\lambda_u = 5 \text{ cm}, N = 89$



Undulator radiation ($K \lesssim 1$)

- Narrow spectral lines
- High spectral brightness
- Partial coherence

$$\lambda = \frac{\lambda_u}{2\gamma^2} \left(1 + \frac{K^2}{2} + \gamma^2 \theta^2 \right)$$

$$K = \frac{eB_0\lambda_u}{2\pi mc}$$

Wiggler radiation ($K \gg 1$)

- Higher photon energies
- Spectral continuum
- Higher photon flux ($2N$)

$$\hbar\omega_c = \frac{3}{2} \frac{\hbar\gamma^2 eB_0}{m}$$

$$n_c = \frac{3K}{4} \left(1 + \frac{K^2}{2} \right)$$



Images: D. Attwood, Soft X-rays and Extreme Ultraviolet Radiation, Cambridge University Press, Cambridge, 1999.

Why undulators and wigglers?

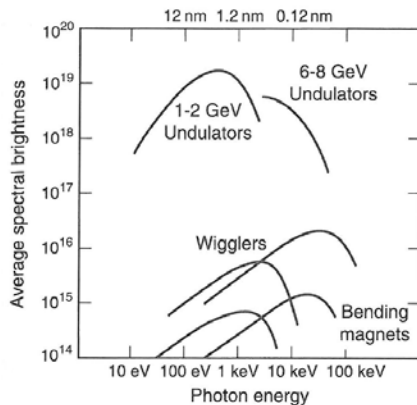
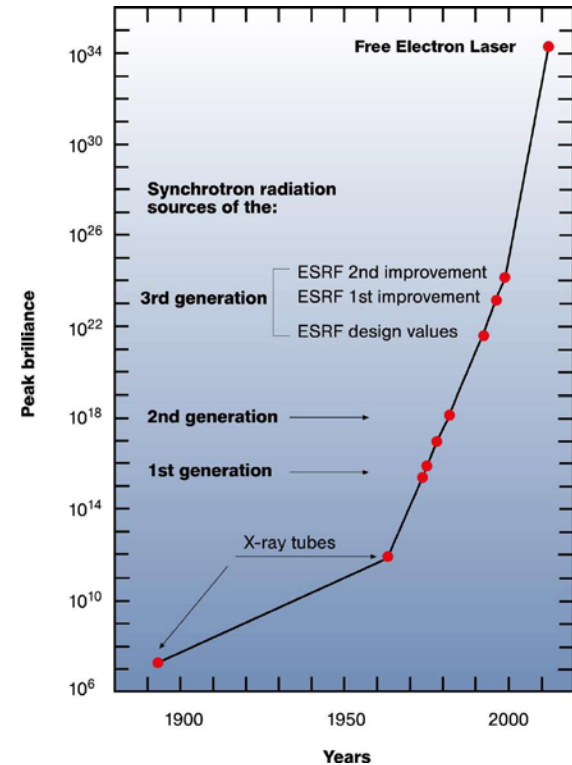


FIGURE 5.24. General trends of spectral brightness for undulator radiation, wiggler radiation, and bending magnet radiation, showing the complementary nature of soft x-ray (1–2 GeV) and hard x-ray (6–8 GeV) storage ring facilities. High spectral brightness is particularly useful for experiments involving scanning microscopy and partial coherence, diffraction from small crystalline samples, and other studies which generally benefit from radiation of minimal divergence emanating from a small source size. Units as in Eq. (5.65).

D. Attwood, Soft X-rays and Extreme Ultraviolet Radiation, Cambridge University Press, Cambridge, 1999.



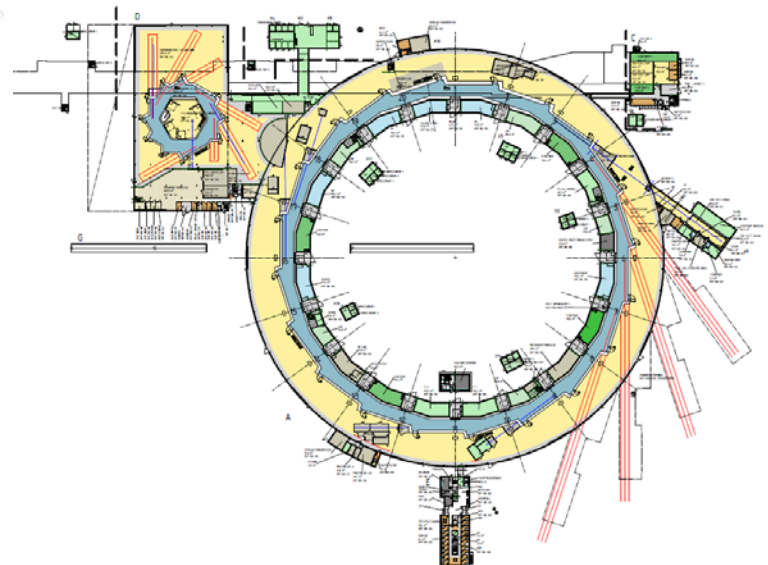
Improved flux / brilliance
as compared to bending magnet sources!

www.hasylab.de



MAX IV – the world's brightest synchrotron radiation source




$$\eta$$

-  Expansionen
-  Vertriebs
-  L&S
-  Kultur
-  Lagerung
-  Techn. Hygiene
-  Technik
-  Technik: Fachgebiet: Vertrieb / K&V / V&M
-  Technik: Fachgebiet: EL
-  Technik: Fachgebiet: V&M
-  Vertrieb
-  Kommunikation
-  K&V
-  Lager

Why will MAX IV be world-best?

Circumference (m)	528
Nr of straight sections	20
Injection	full energy, top-up
Stored current (mA)	500
Horizontal emittance (nm rad)	0.2 - 0.3
Vertical emittance (nm rad)	< 0.008
Horizontal beam size (σ μ m)	42 - 52
Vertical beam size (σ μ m)	< 6

Circumference (m)	96
Nr of straight sections	12
Injection	full energy, top-up
Stored current (mA)	500
Horizontal emittance (nm rad)	6.0
Vertical emittance (nm rad)	0.06
Horizontal beam size (σ μ m)	184
Vertical beam size (σ μ m)	13

Characteristics will allow:
nanofocusing, use of coherence, extremely high resolution,
use of very small biological samples, ...





August 2011



3 GeV ring

Linac (1.5 and 3 GeV)

1.5 GeV ring

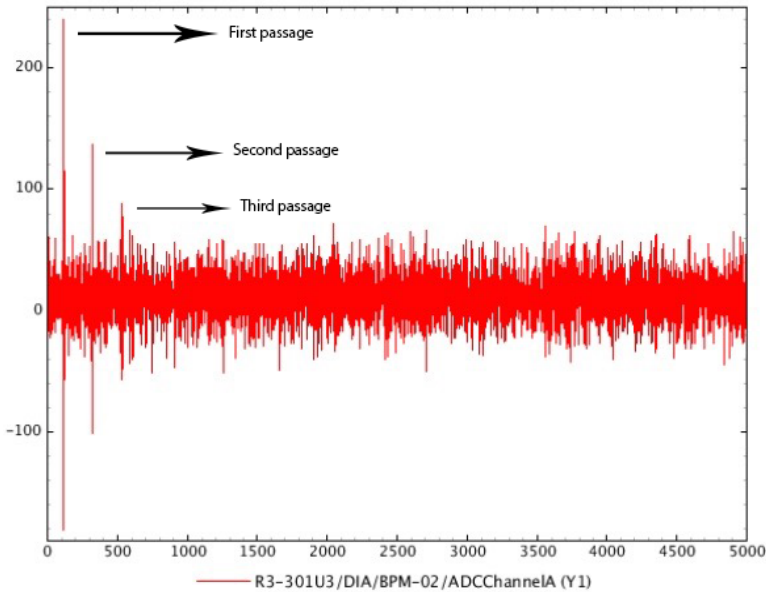


Inauguration, 21 June 2016

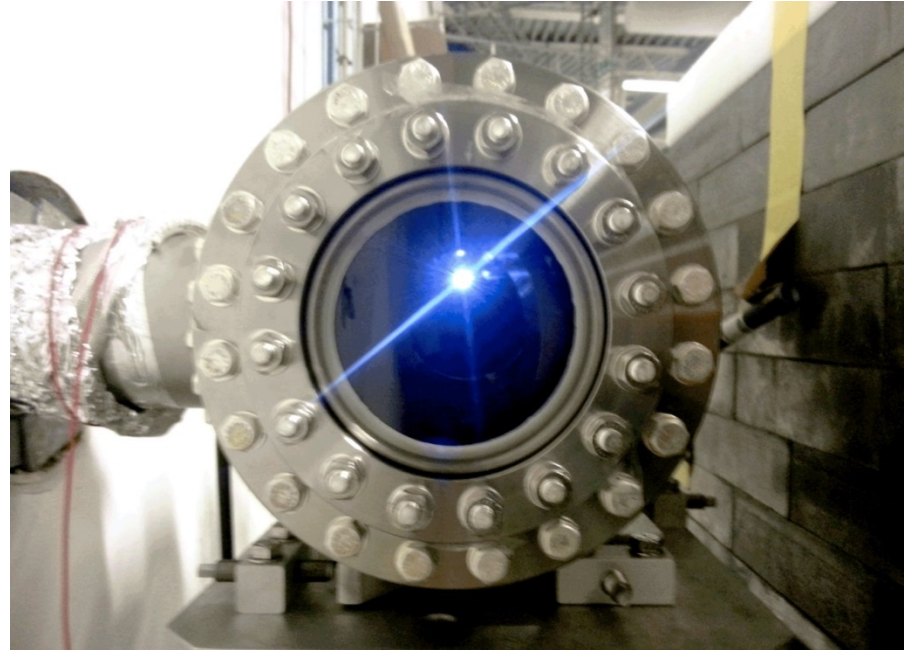
Inauguration: 21 June 2016



Sweden is not dark!



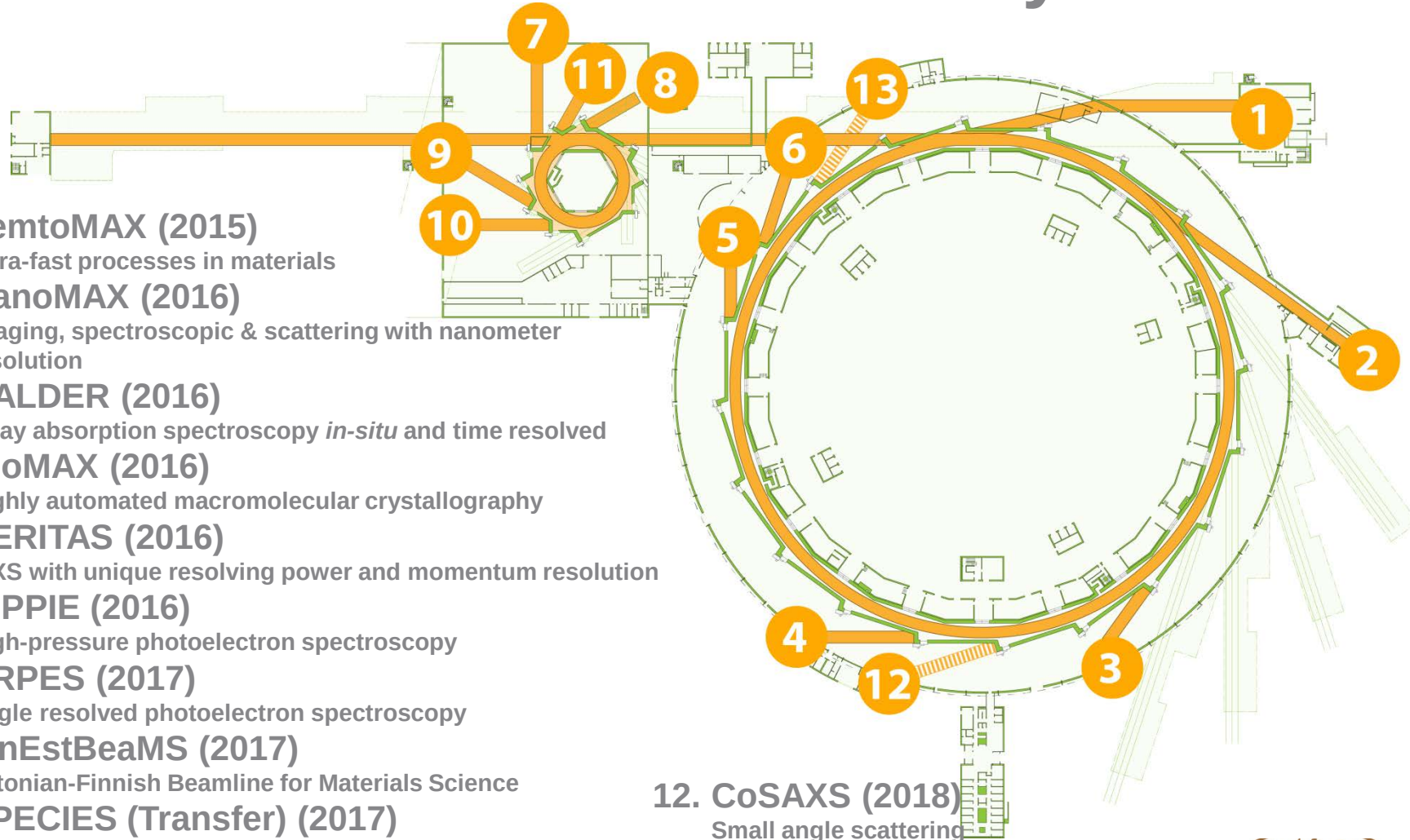
26 August 2015: circulating electron beam
in the 3 GeV ring of MAX IV



2 November 2015: synchrotron radiation
recorded



The MAX IV Laboratory



1. FemtoMAX (2015)

Ultra-fast processes in materials

2. NanoMAX (2016)

Imaging, spectroscopic & scattering with nanometer resolution

3. BALDER (2016)

X-ray absorption spectroscopy *in-situ* and time resolved

4. BioMAX (2016)

Highly automated macromolecular crystallography

5. VERITAS (2016)

RIXS with unique resolving power and momentum resolution

6. HIPPIE (2016)

High-pressure photoelectron spectroscopy

7. ARPES (2017)

Angle resolved photoelectron spectroscopy

8. FinEstBeaMS (2017)

Estonian-Finnish Beamline for Materials Science

9. SPECIES (Transfer) (2017)

VUV High-pressure photoelectron spectroscopy and RIXS

10. FlexPES (Transfer) (2017)

Photoelectron Spectroscopy and NEXAFS

11. MAXPeem (Transfer) (2017)

Photoelectron microscopy

12. CoSAXS (2018)

Small angle scattering

13. SoftiMAX (2018)

Coherent Soft X-Ray Scattering, Holography

14. DanMAX (2019)



Initial beamline programme

- **BioMAX (3 GeV ring):** A multipurpose high throughput beamline for macromolecular crystallography.
- **VERITAS (3 GeV ring):** A beamline for soft X-ray Resonant inelastic X-ray scattering (RIXS) in the energy range of 275-1500 eV.
- **HIPPIE (3 GeV ring):** A state-of-the-art beamline for high pressure X-ray photoelectron spectroscopy (HP-XPS), high pressure X-ray absorption spectroscopy (HP-XAS) as well as XPS and XAS in ultrahigh vacuum.
- **NanoMAX (3 GeV ring):** A hard X-ray beamline for micro- and nanobeams.
- **FemtoMAX (Linac):** A beamline situated on the extension of the linac to facilitate studies of the structure and dynamics of materials with X-ray pulses of 100 fs.
- **ARPES (1.5 GeV ring):** A beamline for angle resolved photo electron spectroscopy (ARPES) covering photon energies between 10-1000 eV but with emphasis on the lower energy range.
- **Balder (3 GeV ring):** A beamline for in-situ hard X-ray spectroscopy.
- **SPECIES (1.5 GeV ring):** A VUV beamline for high pressure X-ray photoelectron spectroscopy (HP-XPS) and Resonant inelastic X-ray scattering (RIXS)
- **FinEstBeaMS (1.5 GeV ring):** A beamline for soft x-ray spectroscopy on vapours and materials
- **SoftiMAX (3 GeV ring):** A beamline for coherent x-ray imaging and for scanning transmission x-ray microscopy (STXM).
- **CoSAXS (3 GeV ring):** A beamline for small angle x-ray scattering (SAXS).
- **FlexPES/PEEM (1.5 GeV ring):** Transfer of the Photoemission electron microscopy and of a soft x-ray spectroscopy beamline from present MAX-lab.



Applications of synchrotron radiation



Applications of synchrotron radiation

Spectroscopy

X-ray absorption spectroscopy, including
X-ray magnetic circular dichroism
X-ray emission spectroscopy
Photoelectron spectroscopy, including
Angle-resolved photoemission spectroscopy
Vibrational spectroscopy

...

Imaging

X-ray tomography
Synchrotron infrared microspectroscopy
Photoemission electron microscopy
Scanning x-ray microscopy
Phase contrast microscopy

...

Scattering

Powder diffraction, crystallography
Small angle x-ray scattering
Inelastic x-ray scattering
Magnetic scattering
Time resolved x-ray scattering
...

Microfabrication

X-ray lithography

List by no means complete!



Applications of synchrotron radiation

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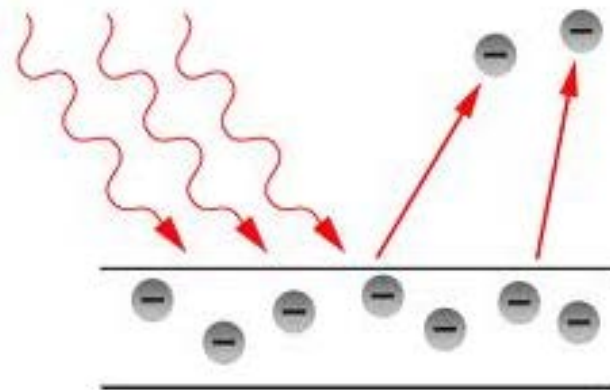
X-ray photoelectron spectroscopy



(X-ray) Photoelectron spectroscopy

Photon in – electron out, i.e. PES is an *electron spectroscopy*

Photoelectric effect (observed by Heinrich Herz 1887, explained by Albert Einstein in 1905)



www.physicsforum.com

Works (of course) on atoms, molecules, and solids

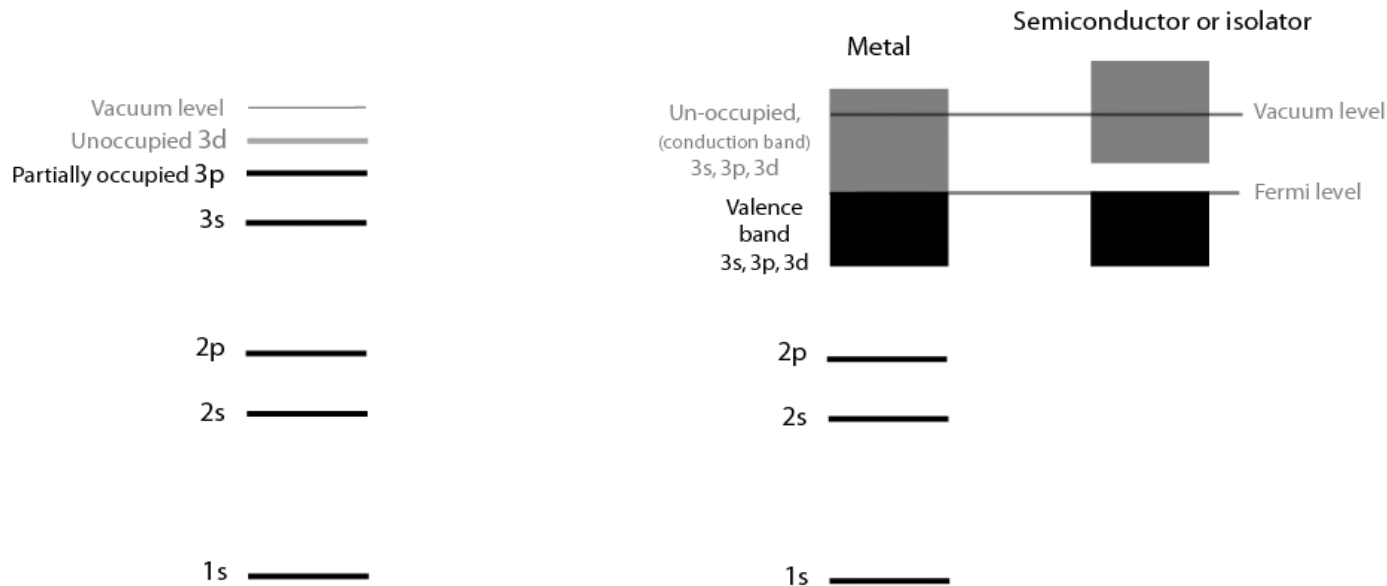


Energy levels of atoms and solids

All electron spectroscopy methods rely on the electronic structure of atoms, molecules, and solids

Schematic energy level diagram for a solid

Schematic energy level diagram for an atom



Note: Energies not to scale



X-ray photoelectron spectroscopy

Photoelectron spectroscopy = Photoemission spectroscopy

XPS = X-ray photoelectron spectroscopy

UPS = Ultraviolet photoelectron spectroscopy

Hard x-ray:

approx. 1200 – 250,000 eV

Soft x-rays:

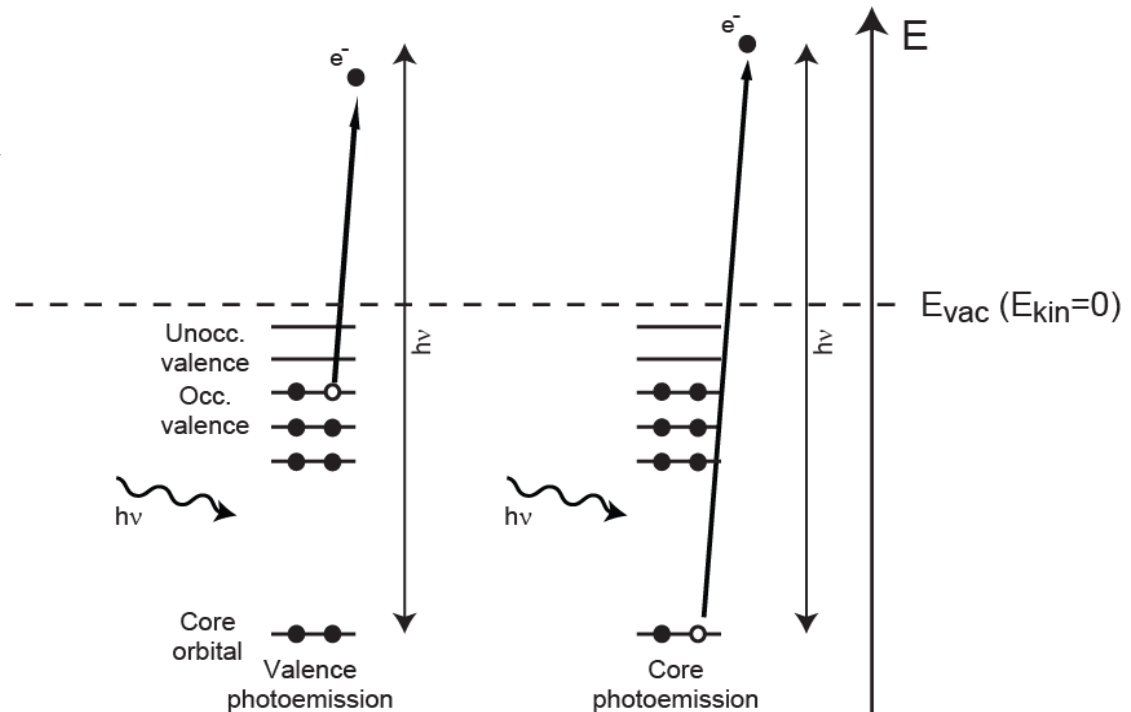
approx. 20 – 1200 eV

Vacuum ultraviolet:

approx. 6 – 20 eV

Ultraviolet:

approx. 3 – 6 eV



X-ray photoelectron spectroscopy

is given by
experiment settings
(tunable at synchrotron
radiation facility!) can be determined

$$h\nu = E_B + E_{\text{kin}} + \phi$$

quantity of
interest!

can be measured

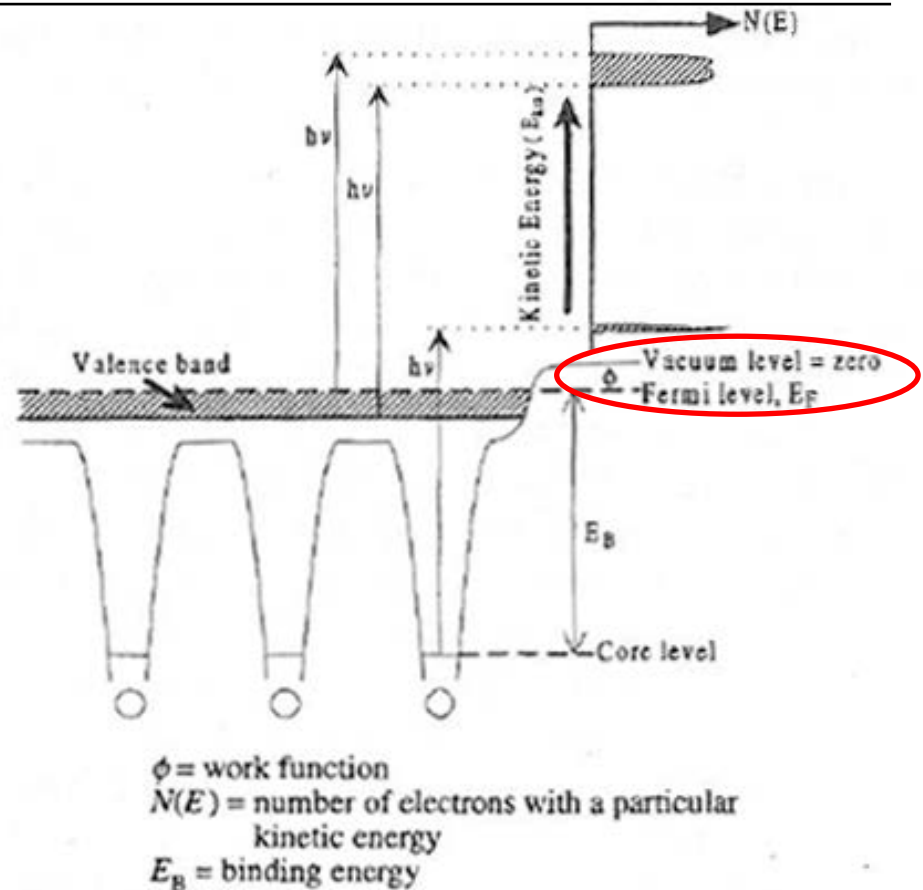


Fig. 2.2 The energetics of an X-ray photoemission experiment.



Core level binding
energies:

characteristic for
the elements

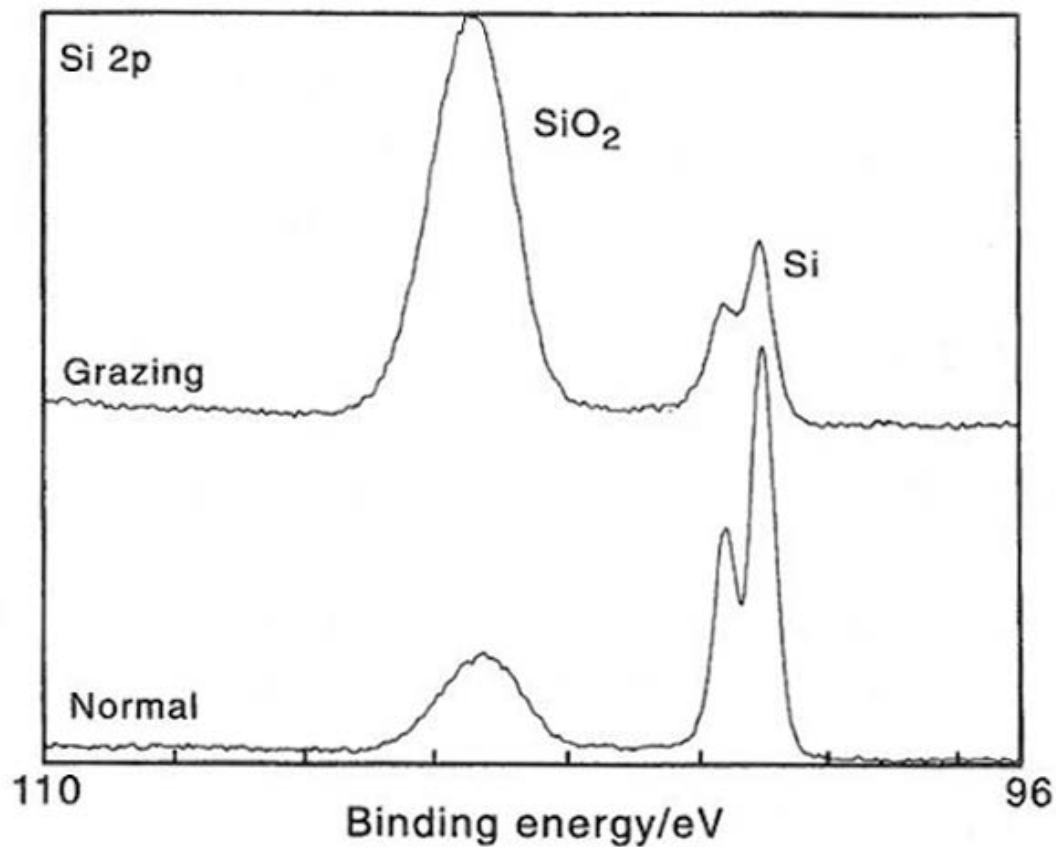
| Binding energies

	1s	2s	2p1	2p3	3s	3p1	3p3	3d3	3d5	4s	4p1	4p3	4d3	4d5	4f5	4f7	5s	5p1	5p3		
1 H	14																			H	55 Cs
2 He	25																			He	56 Ba
3 Li	55																			Li	57 La
4 Be	112																			Be	58 Ce
5 B	188																			B	59 Pr
6 C	284																			C	60 Nd
7 N	410	2s	2p1	2p3	3s	3p1	3p3	3d3	3d5	4s	4p1	4p3	4d3	4d5	4f5	4f7	5s	5p1	5p3	N	61 Pm
8 O	543	23																		O	62 Sm
9 F	686	30																		F	63 Eu
10 Ne	863	41	14	14																Ne	64 Gd
11 Na	1072	64	31	31																Na	65 Tb
12 Mg		90	51	51																Mg	66 Dy
13 Al		119	74	74																Al	67 Ho
14 Si		153	103	102																Si	68 Er
15 P		191	134	133	14															P	69 Tm
16 S		229	166	165	17															S	70 Yb
17 Cl		270	201	199	17															Cl	71 Lu
18 Ar		319	243	241	22	3p1	3p3	3d3	3d5	4s	4p1	4p3	4d3	4d5	4f5	4f7	5s	5p1	5p3	Ar	72 Hf
19 K		378	296	293	33	17	17													K	73 Ta
20 Ca		439	350	347	44	25	25													Ca	74 W
21 Sc		501	407	402	53	31	31													Sc	75 Re
22 Ti		565	464	458	62	37	37													Ti	76 Os
23 V		630	523	515	69	40	40													V	77 Ir
24 Cr		698	586	577	77	46	45													Cr	78 Pt
25 Mn		770	652	641	83	49	48													Mn	79 Au
26 Fe		847	723	710	93	56	55													Fe	80 Hg
27 Co		927	796	781	103	63	61													Co	81 Ti
28 Ni		1009	873	855	112	69	67													Ni	82 Pb
29 Cu		1098	954	934	124	79	77													Cu	83 Bi
30 Zn		1196	1045	1022	140	92	89	10	10											Zn	84 Po
31 Ga			1144	1117	160	108	105	20	20											Ga	85 At
32 Ge					184	128	124	32	31											Ge	86 Rn
33 As					207	148	143	45	44											As	87 Fr
34 Se					232	169	163	58	57											Se	88 Ra
35 Br					256	189	182	70	69	4s	4p1	4p3	4d3	4d5	4f5	4f7	5s	5p1	5p3	Br	89 Ac
36 Kr					287	216	208	89	88	22										Kr	90 Th
37 Rb					322	247	238	111	110	29	14	14								Rb	91 Pa
38 Sr					358	280	269	135	133	37	20	20								Sr	92 U
39 Y					395	313	301	160	158	45	25	25								Y	93 Np
40 Zr					431	345	331	183	181	51	29	29								Zr	94 Pu
41 Nb					470	379	364	209	206	59	35	35								Nb	95 Am
42 Mo					508	413	396	233	230	65	38	38								Mo	96 Cm
43 Tc					544	445	425	257	253	68	39	39								Tc	
44 Ru					587	485	463	286	282	77	45	45								Ru	
45 Rh					629	522	498	314	309	83	49	49								Rh	
46 Pd					673	561	534	342	337	88	54	54								Pd	
47 Ag					718	604	573	374	368	97	58	58	4d3	4d5	4f5	4f7	5s	5p1	5p3	Ag	
48 Cd					772	652	618	412	405	109	68	68	11	11						Cd	
49 In					828	704	666	453	445	123	79	79	19	19						In	
50 Sn					884	757	715	494	486	137	91	91	26	25						Sn	
51 Sb					946	814	768	539	530	155	105	105	35	34						Sb	
52 Te					1009	873	822	585	575	171	114	114	44	43			14			Te	
53 I					1071	930	874	630	619	186	123	123	52	50			16			I	
54 Xe					1144	997	936	685	672	209	141	141	65	63			19			Xe	

1064	997	738	724	230	170	158	77	75	24	Cs			
1137	1062	795	780	254	192	179	92	90	23	Ba			
1126	851	834	274	210	195	104	101		34	17	17	La	
1184	900	882	290	222	207	112	108		37	18	18	Ce	
	950	930	305	237	218	114	114		38	20	20	Pr	
	1001	980	318	248	227	120	120		38	23	23	Nd	
	1060	1034	337	264	242	129	129		38	22	22	Pm	
	1110	1083	349	283	250	132	132		41	20	20	Sm	
	1166	1136	366	289	261	136	136		34	24	24	Eu	
		1186	380	301	270	141	141		36	21	21	Gd	
			398	317	284	150	150		42	28	28	Tb	
			412	329	293	154	154		63	26	26	Dy	
			431	345	306	161	161		51	20	20	Ho	
			451	362	320	169	169		61	25	25	Er	
			470	378	333	180	180		54	32	26	Tm	
			483	392	342	194	185		55	33	26	Yb	
			507	412	359	207	197		58	34	27	Lu	
			537	437	382	224	213	19	17	64	37	Hf	
			566	464	403	241	229	27	25	71	45	37	Ta
			594	491	425	257	245	36	34	77	47	37	W
			628	521	449	277	263	45	43	81	44	33	Re
			657	549	475	294	279	55	52	86	60	48	Os
			692	579	497	313	297	65	62	98	65	53	Ir
			726	610	521	333	316	76	73	105	69	54	Pt
			763	643	547	354	336	89	85	110	75	57	Au
			803	681	577	379	359	104	100	127	84	65	Hg
			845	721	608	406	385	122	118	137	100	76	Tl
			893	762	645	435	413	143	138	148	107	84	Pb
			942	807	681	467	443	164	159	161	120	94	Bi
										184			Po
										210			At
										238			Rn
										268			Fr
										299			Ra
										319			Ac
										333			Th
	1168	968	714	677	344	335	290	225	179	Pa			
		1046	781	739	391	380	325	262	197	U			
			1086	816	771	414	402		206	Np			
			1121	850	802	439	427		216	Pu			
				883	832	463	449		216	Am			
				919	865	487	473		232	Cm			



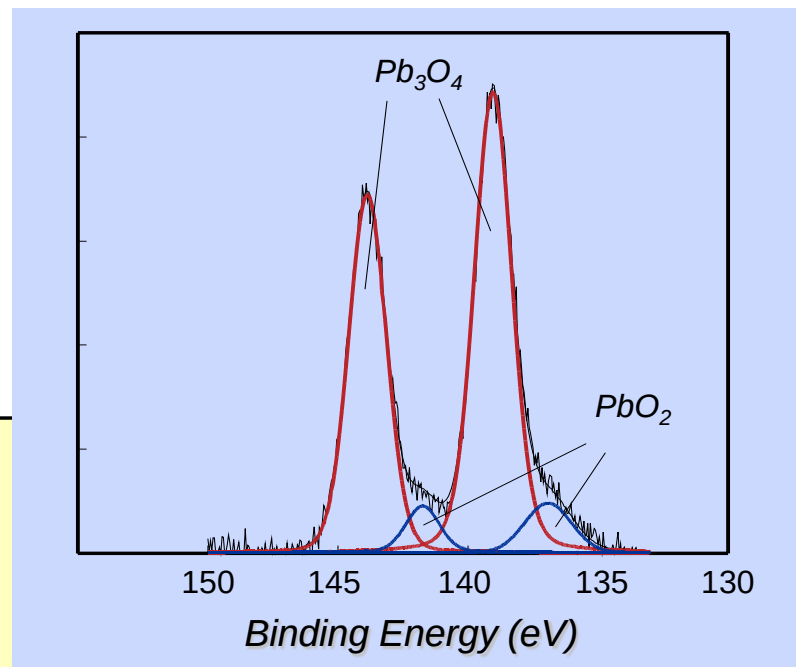
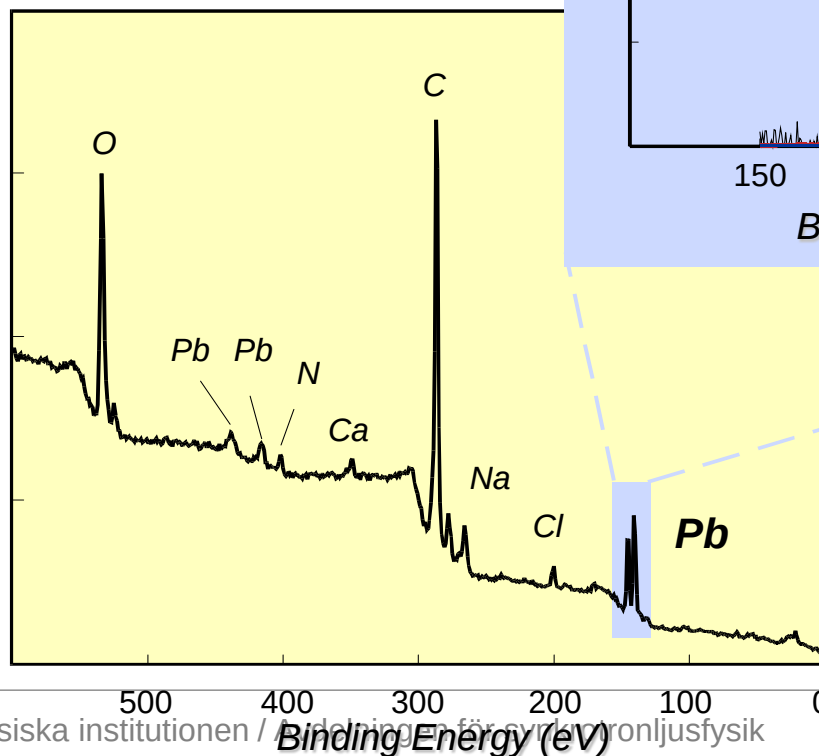
Chemical shifts in x-ray photoelectron spectroscopy



XPS Analysis of Pigment from Mummy Artwork



*Egyptian Mummy
2nd Century AD
World Heritage Museum
University of Illinois*

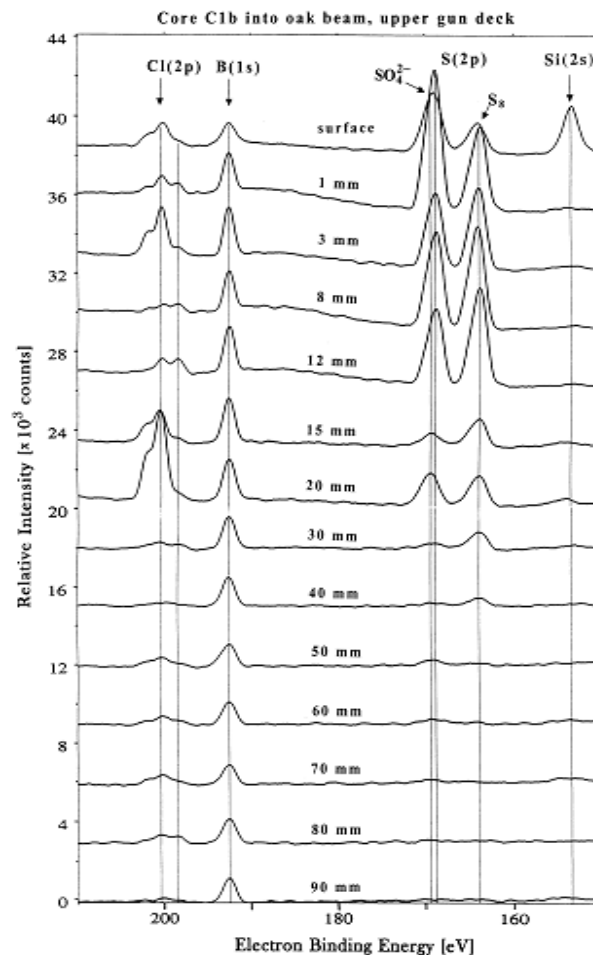


*XPS analysis showed
that the pigment used
on the mummy
wrapping was Pb_3O_4
rather than PbO_2*

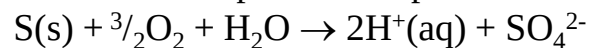


Saving the Vasa with XPS?

http://www-ssrl.slac.stanford.edu/research/highlights_archive/vasa.html



The study shows that in humid museum atmospheres a stepwise sulfur oxidation produces sulfenic acid:



Mean free path I

XPS and AES rely on the **short mean free path of low energy electrons** in solids for achieving **surface sensitivity**.

The intensity removed ($-dI$) per length travelled (dx)

$$-dI = \sigma N' I dx$$

(σ : cross section for inelastic processes)

(σ : cross section for inelastic processes)

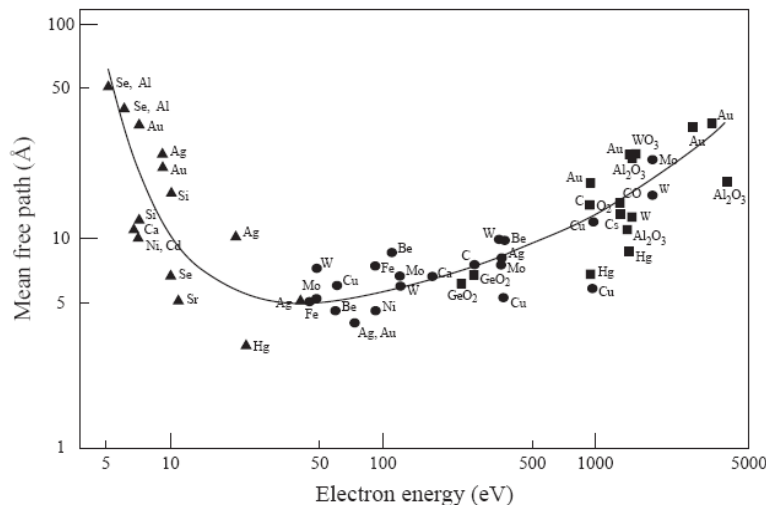
(N' : Scattering centers per cm^3)

$$I(x) = I_0 e^{-\sigma N' x} = I_0 e^{-x/\lambda}$$

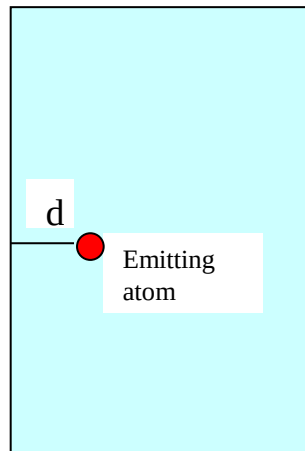
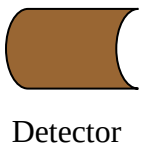
where $\lambda = (\sigma N')^{-1}$ is the mean free path

$I(x)$ is the intensity of electrons that have **not** lost any energy after they have travelled the distance x in the solid.

So, if you made all atoms in a solid emit electrons at a given energy of around say 70 eV and detected all electrons coming out of the sample **with that energy**, the majority of the electrons would come from the first few atomic layers.

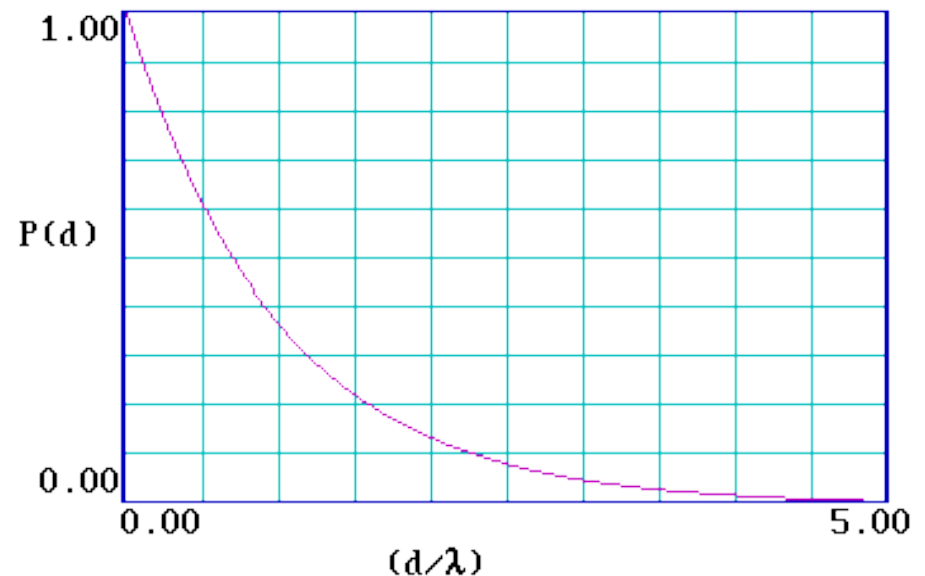


Mean free path I



Probability of an electron travelling the distance d through a material without losing energy (λ : mean free path)

$$P(d) = e^{-d/\lambda} \quad (\text{remember } \lambda = \lambda(E))$$



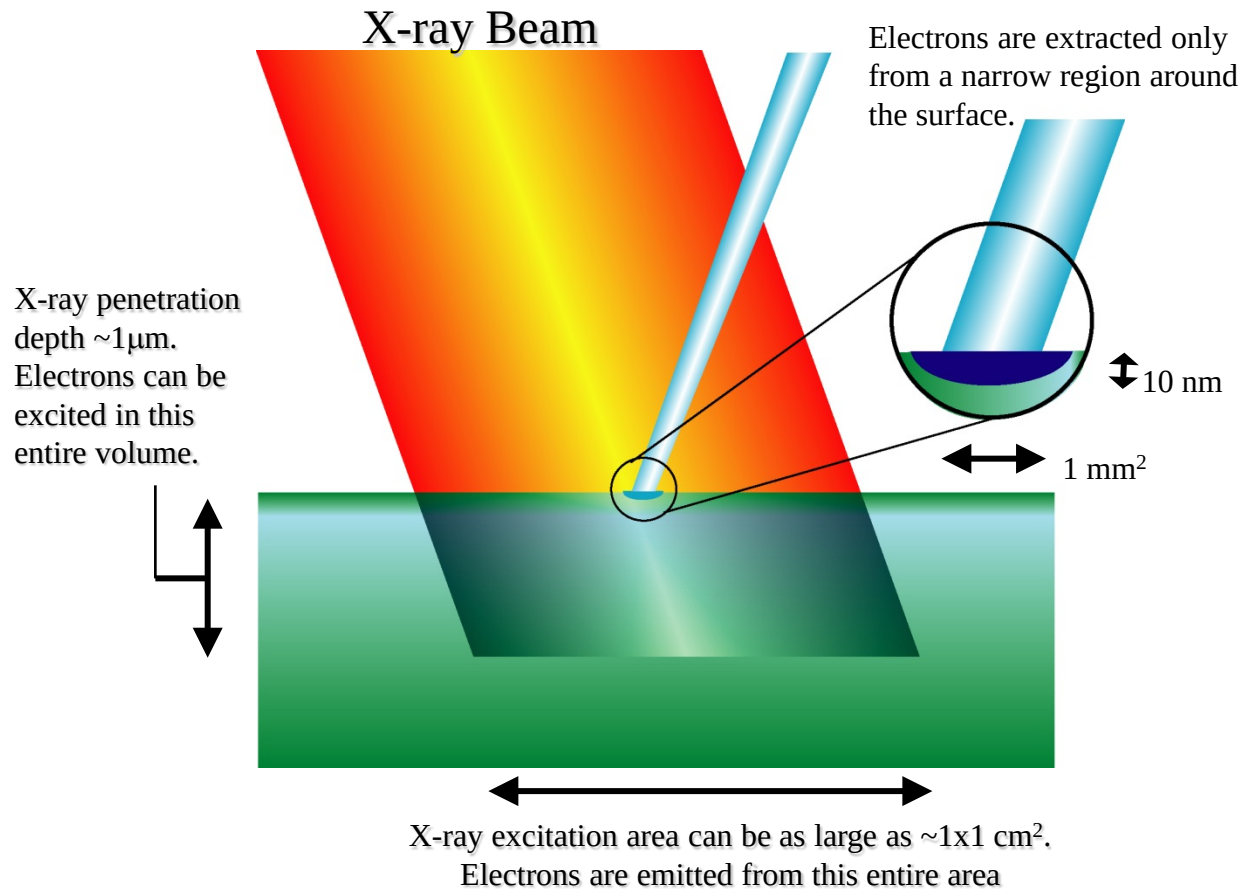
As you see virtually no electrons make it for more than 5λ without losing energy.

Actually most of the electrons which escape from a surface without losing energy have originated from within $1-2 * \lambda$ below the surface. Remember the minimum λ is about 5 Å.



X-ray photoelectron spectroscopy

Photon energies: 100-2000eV, Electron energies: 0-1000eV



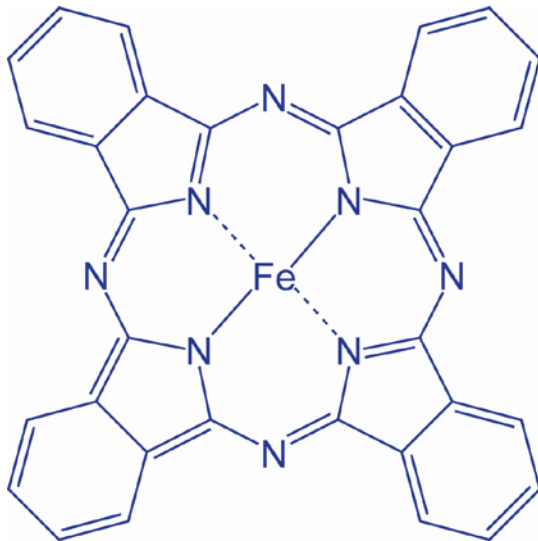
X-ray photoelectron spectroscopy

- **delivers elemental information**
- **delivers chemical information**
- **on solids: very surface sensitive**



Iron phthalocyanine

Iron phthalocyanine
(FePc)



- Similar to haem (responsible for oxygen storage and transport in mammals)
- Iron in ionic state (+2)
- Iron behaves very much as a single atom, but is modified due to presence of "macrocycle"



What are the electron configurations of Fe and Fe(II)? What are the spin quantum numbers of Fe and Fe(II)?

Do not confuse with notation from atomic spectroscopy!

Atomic spectroscopy: Fe I = neutral helium, Fe II = Fe^+ , etc.

Chemical notation: Fe(I) = Fe^+ , Fe(II) = Fe^{2+} , etc.

IA

1

H

IIA

4

Be

3

Li

11

Na

19

K

37

Rb

55

Cs

87

Fr

2

He

10

Ne

18

Ar

36

Kr

54

Xe

86

Rn

5

B

13

Al

31

Ga

49

In

81

Tl

6

C

14

Si

32

Ge

50

Sn

82

Pb

7

N

15

P

33

As

51

Sb

83

Bi

8

O

16

S

34

Se

52

Te

84

Po

9

F

17

Cl

35

Br

53

I

85

At

21

Sc

22

Ti

23

V

24

Cr

25

Mn

26

Fe

27

Co

28

Ni

29

Cu

30

Zn

40

Zr

41

Nb

42

Mo

43

Tc

44

Ru

45

Rh

46

Pd

47

Ag

48

Cd

72

Hf

73

Ta

74

W

75

Re

76

Os

77

Ir

78

Pt

79

Au

80

Hg

104

Rf

105

Ha

106

Sg

107

Ns

108

Hs

109

Mt

110

110

111

111

112

112

113

113

III B

IV B

V B

VI B

VII B

— VII —

IB

IB

III A

IV A

V A

VI A

VII A

0

Periodic Table

of the Elements

* Lanthanide
Series

+ Actinide
Series

58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu
90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr



What are the electron configurations of Fe and Fe(II)?

What are the spin quantum numbers of Fe and Fe(II)?

Hund's rules for ground state configurations:

- (a) Highest possible S
- (b) Highest possible L for the S from (a)
- (c) Highest possible J for more than half-filled shell, smallest possible J for less than half-filled shell

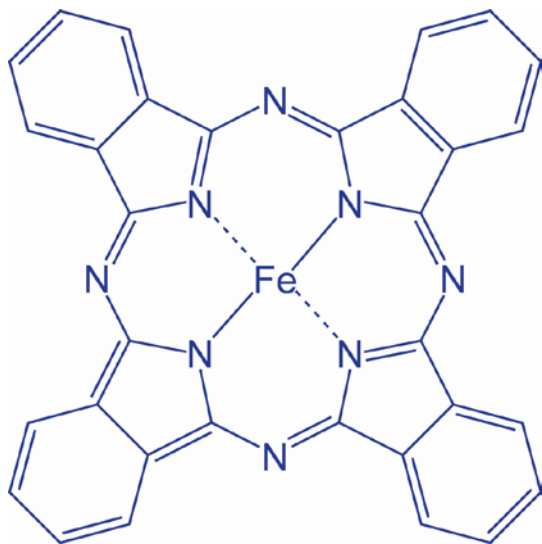


What are the electron configurations of Fe and Fe(II)?
What are the spin quantum numbers of Fe and Fe(II)?



What are the electron configurations of Fe and Fe(II)?

What are the spin quantum numbers of Fe and Fe(II)?



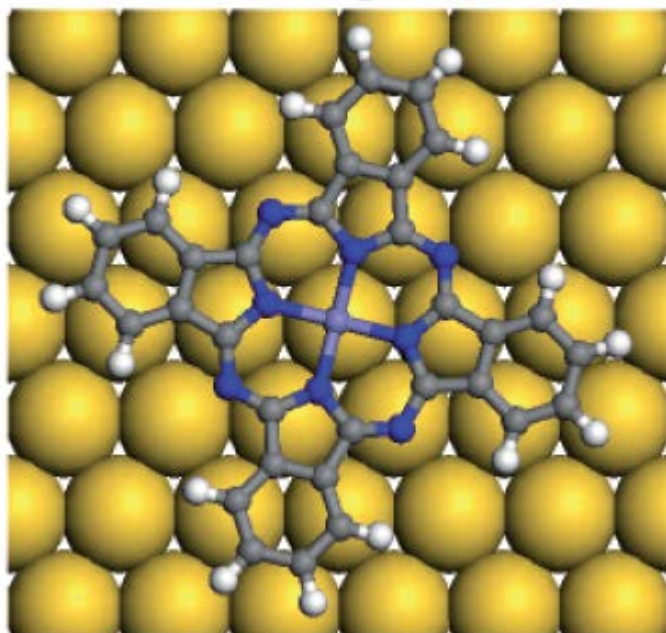
In general, in transition metal complexes the electron configurations are hydrogen-like!

Furthermore the d levels split up due to the "ligand field".

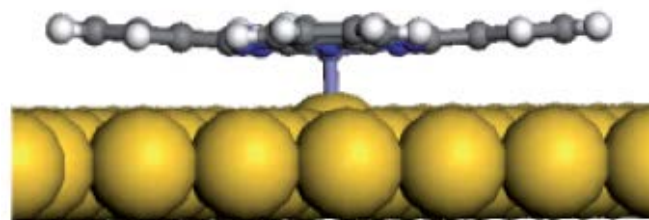


Iron phthalocyanines on a Au(111) surface

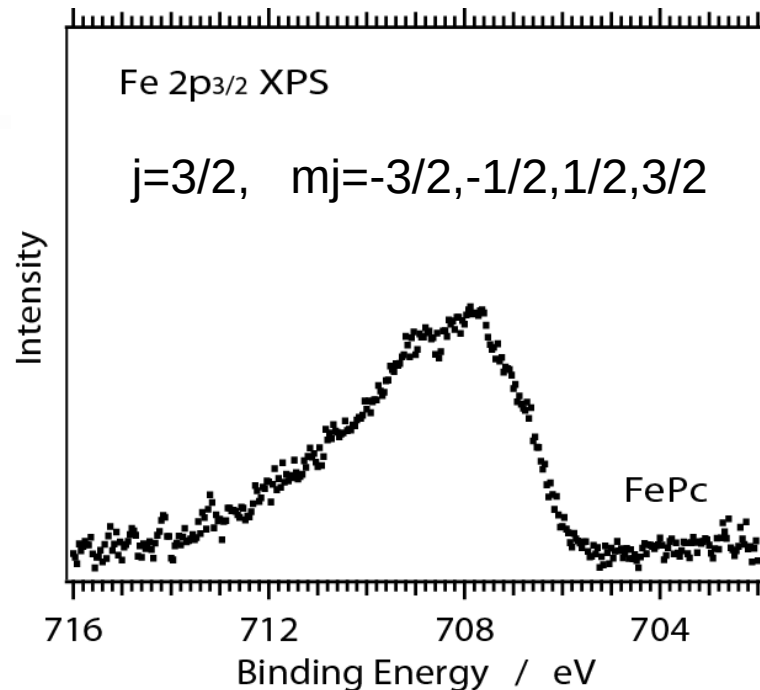
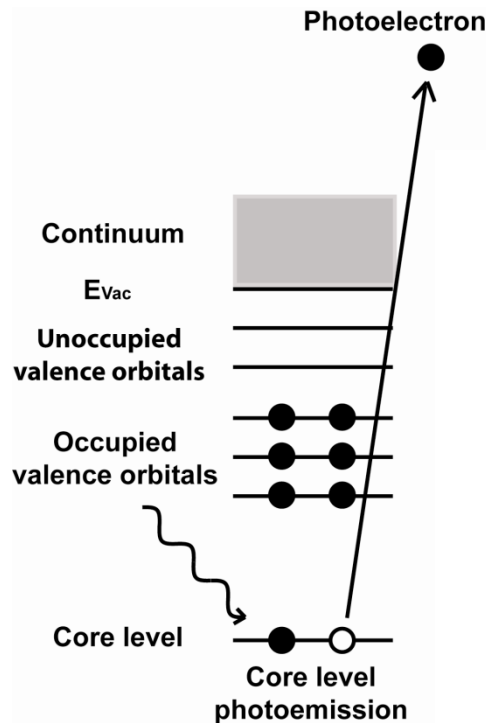
b) Configuration 2



f) Configuration 2



X-ray photoelectron spectroscopy



We just said that $S=1$ and further L needs to be an integer number. How can J be a half-integer?

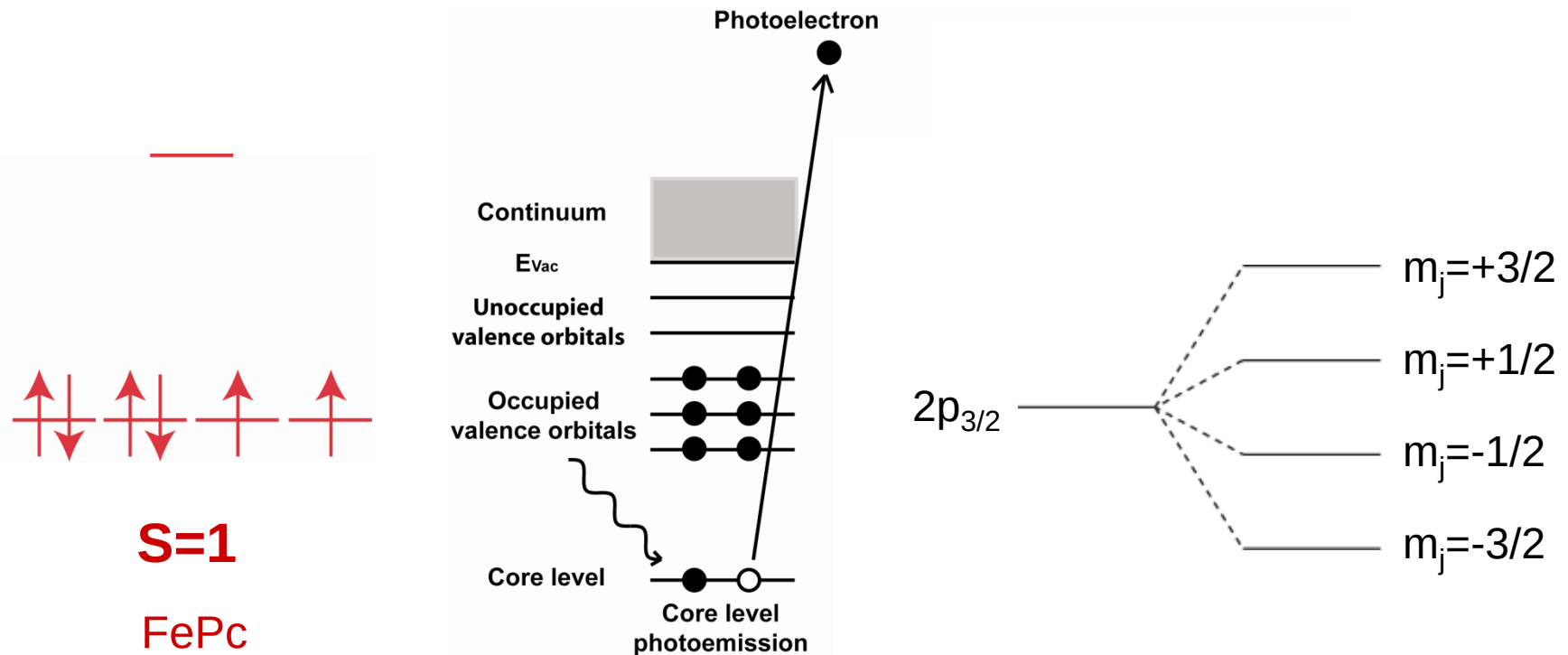
" $2p_{3/2}$ " is about the angular momentum of the core (hole) level only!



X-ray photoelectron spectroscopy



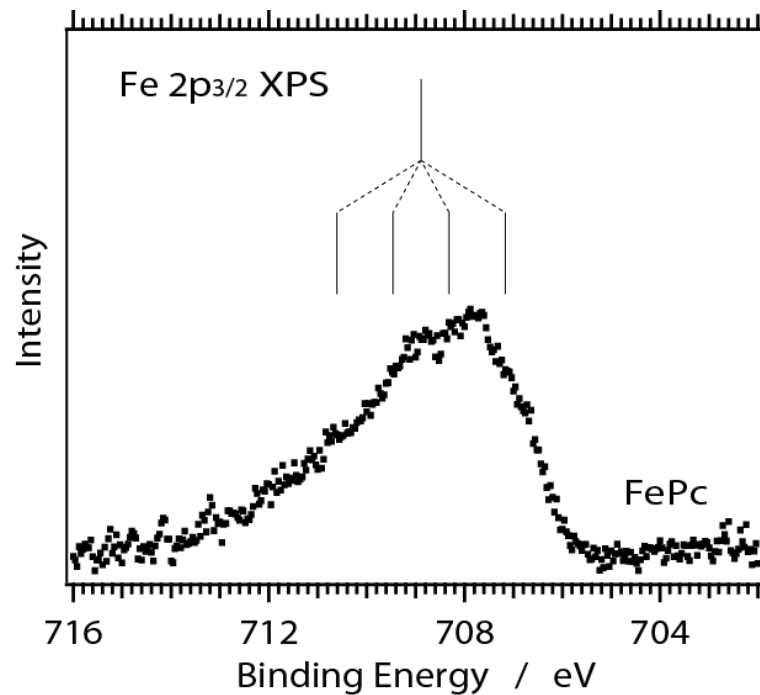
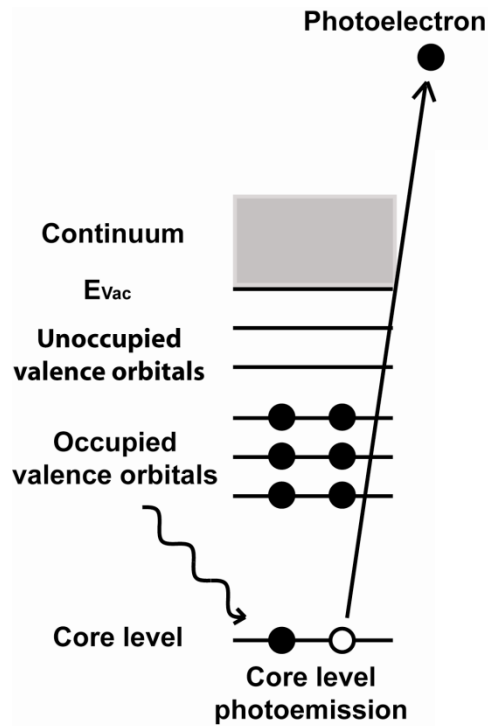
Metal valence levels of iron phthalocyanine



Coupling of valence spin angular momentum
with angular momentum of the core hole!



X-ray photoelectron spectroscopy

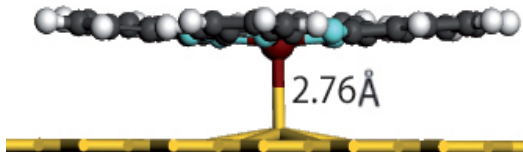


$$j=3/2, \quad m_j=-3/2, -1/2, 1/2, 3/2$$



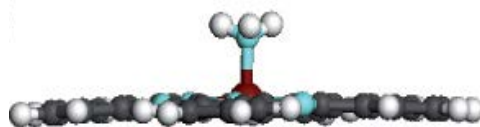
Pyridine, carbon monoxide, nitric oxide on iron phthalocyanine

FePc/Au(111)



2.76 Å

NH₃/FePc

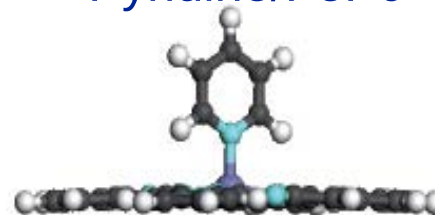


4.04 Å

Binding energy: 0.55 eV

$$\mu_s = 0.27 \mu_B$$

Pyridine/FePc



4.14 Å

Binding energy: 0.6 eV

$$\mu_s = 0.25 \mu_B$$

CO/FePc



4.18 Å

Binding energy: 1.4 eV

$$\mu_s = 0.09 \mu_B$$

NO/FePc



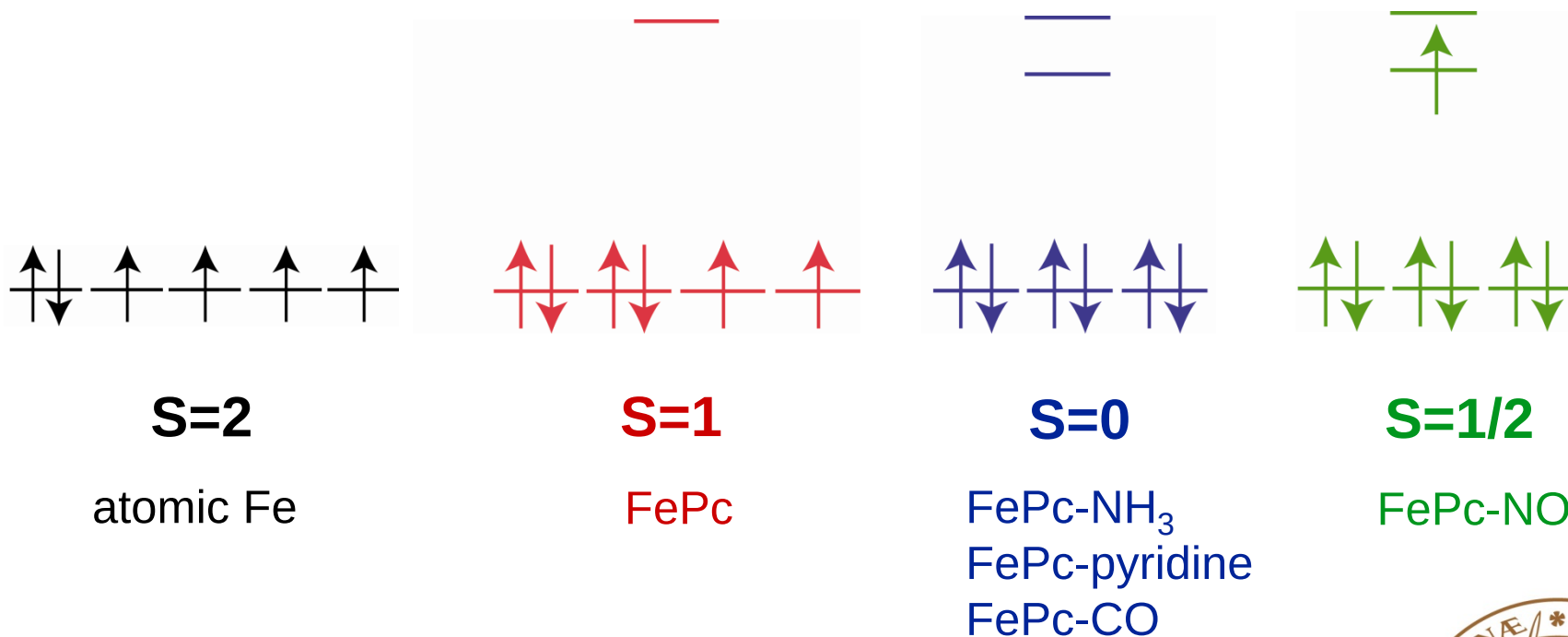
4.34 Å

Binding energy: 1.6 eV

$$\mu_s = 0.84 \mu_B$$

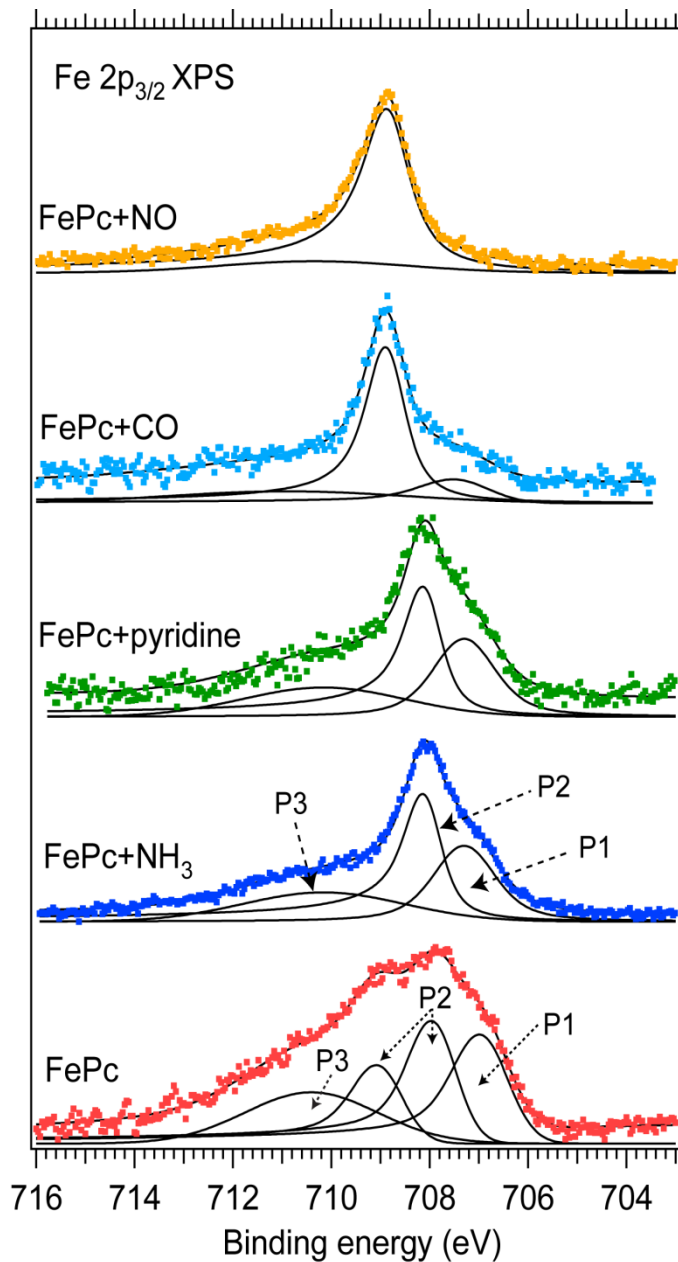


d-levels of iron phthaloycanine



Ligand field splitting





	Width (eV)	S
FePc/Au(111)	2.29	1
NH ₃ /FePc/Au(111)	1.02	0
Py/FePc/Au(111)	1.02	0
CO/FePc/Au(111)	1.01	0
NO/FePc/Au(111)	1.26	1/2

NH₃, pyridine, CO quench the spin; NO reduces the spin



X-ray absorption spectroscopy

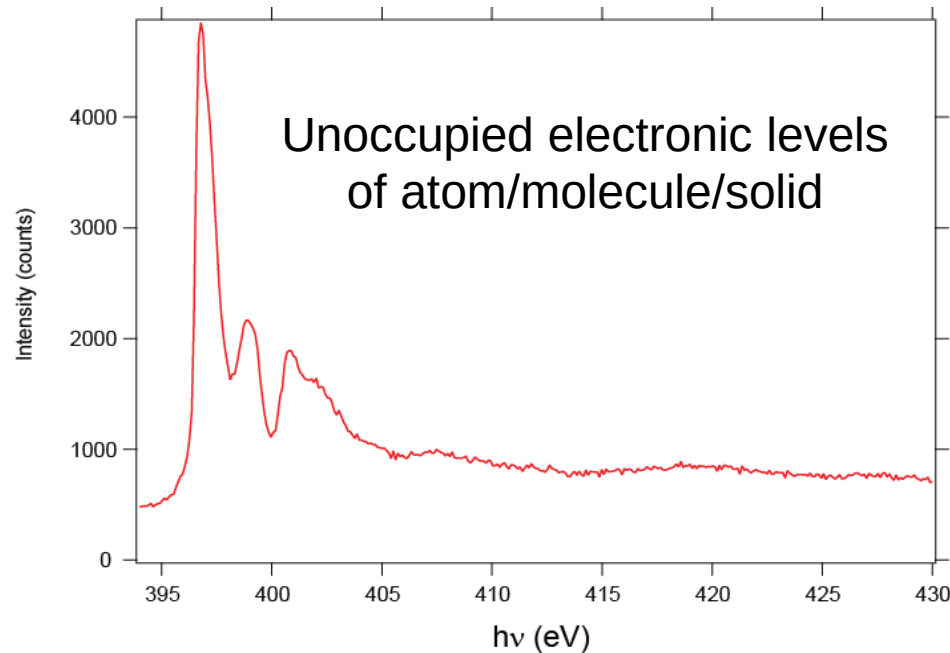
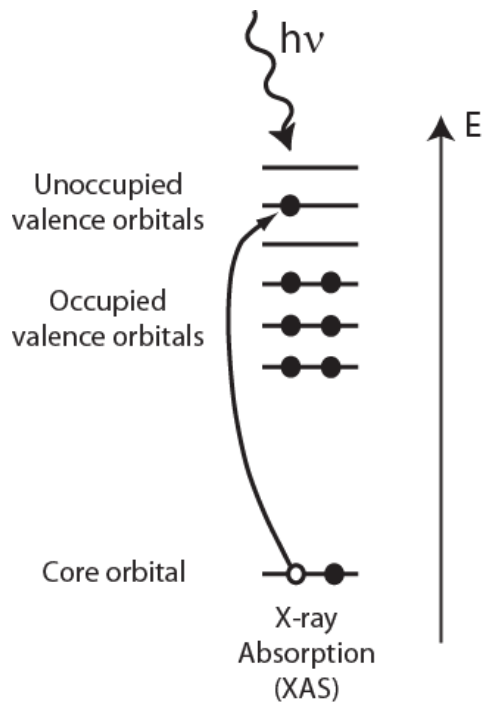


X-ray absorption spectroscopy

XANES = X-ray Absorption Near Edge Structure

NEXAFS = Near Edge X-ray Absorption Fine Structure

XAS = X-ray Absorption Spectroscopy



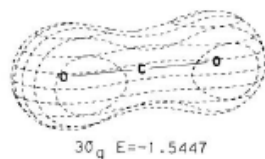
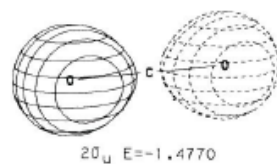
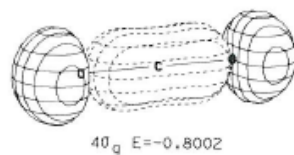
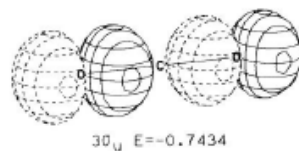
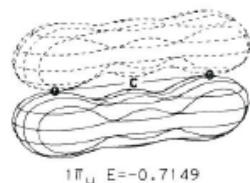
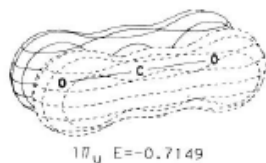
Linear combination of atomic orbitals



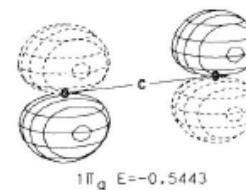
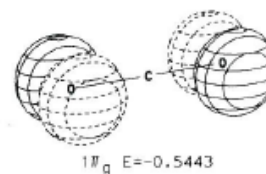
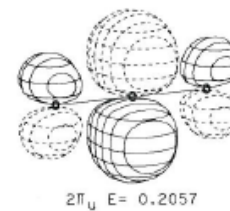
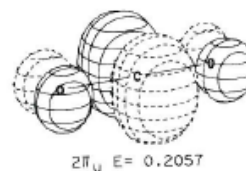
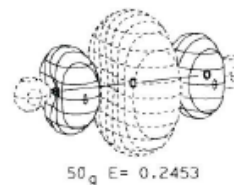
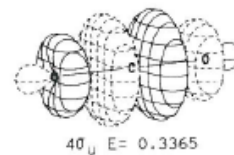
σ^* and π^* orbitals

45. Carbon Dioxide

Symmetry: $D_{\infty h}$



Carbon Dioxide (Continued)



Excitations in x-ray absorption

Electron removed typically from 1s orbital

Dipole selection rule: $\Delta l=1$

Implies that electron only can be put into *atomic* orbital with $l=1$ (p) only

Both σ^* and π^* orbitals have p atomic orbital character –

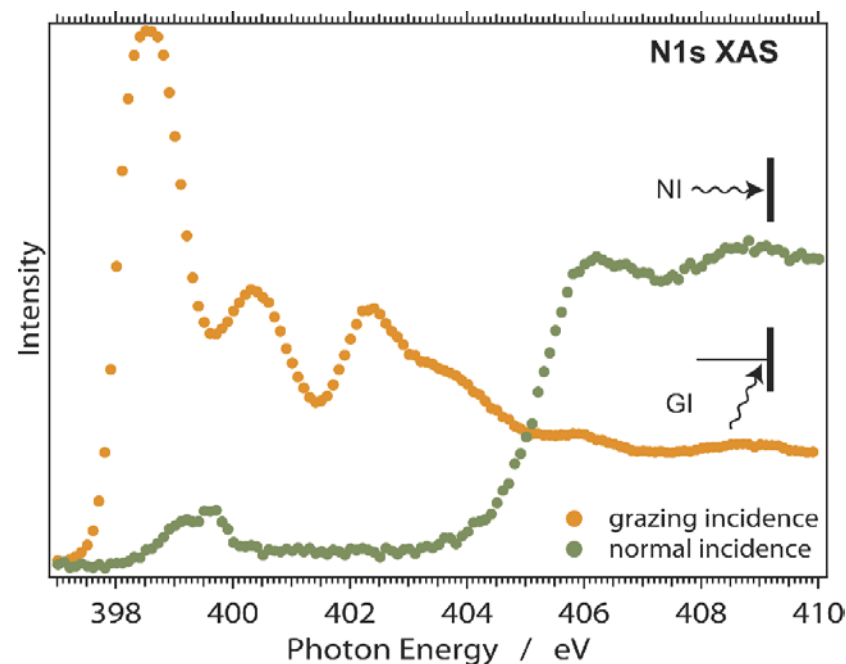
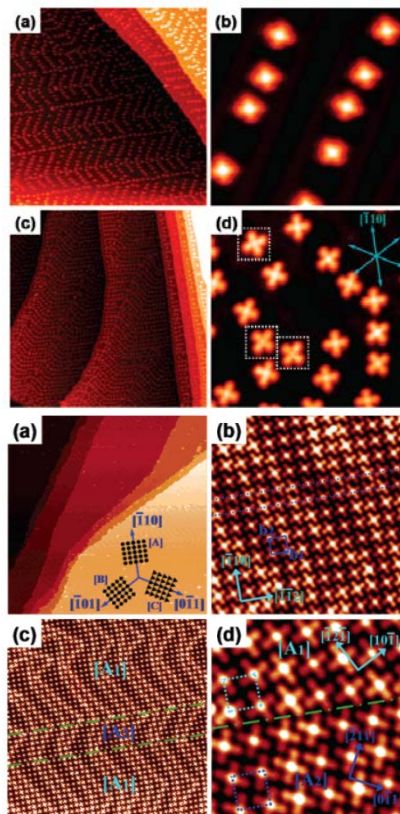
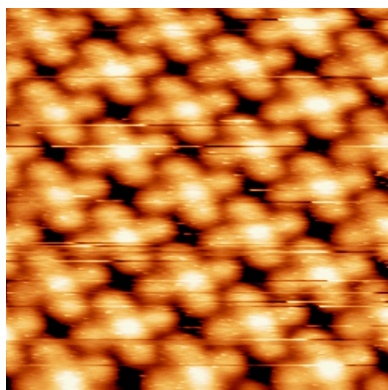
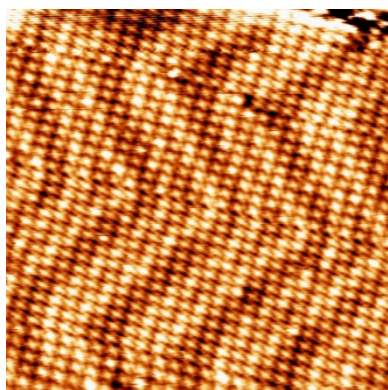
σ^* along bond axis

π^* perpendicular to bond axis

→ **makes possible determination of molecular geometry
at surface**



Geometry determination for iron phthalocyanine on a Au(111) surface



Z. H. Cheng et al.,
J. Phys. Chem. C 111 (2007) 9240
J. Phys. Chem. C 111 (2007) 2656



Take-home messages

Synchrotron radiation:

- Extremely powerful tool for investigation of matter (atoms, molecules, solids)
- Extremely high "brilliance" (photon flux into solid angle)
- Lunds hosts a world-leading synchrotron radiation facility – MAX IV

X-ray photoelectron spectroscopy:

- Photon in – electron out
- Probes occupied states
- Elemental analysis
- Chemical analysis (down to spin)
- when applied to solids: very surface sensitive

X-ray absorption spectroscopy:

- Probes unoccupied states
- Chemical analysis
- Geometry determination

